



SHERPA

CIAO's Modeling and Fitting Application

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Modeling and Fitting Software

- **XSPEC** - analysis of 1D X-ray data (imaging + grating)
- **ISIS** and **Pint of Ale**- primarily for analysis of high-resolution (ie grating) X-ray data
- **Sherpa** - generalised multi-dimensional fitting package
- All programs use the technique of **forward fitting**:
 - a model is evaluated, compared to the actual data, and then the parameters are changed to improve the match. This is repeated until convergence occurs.



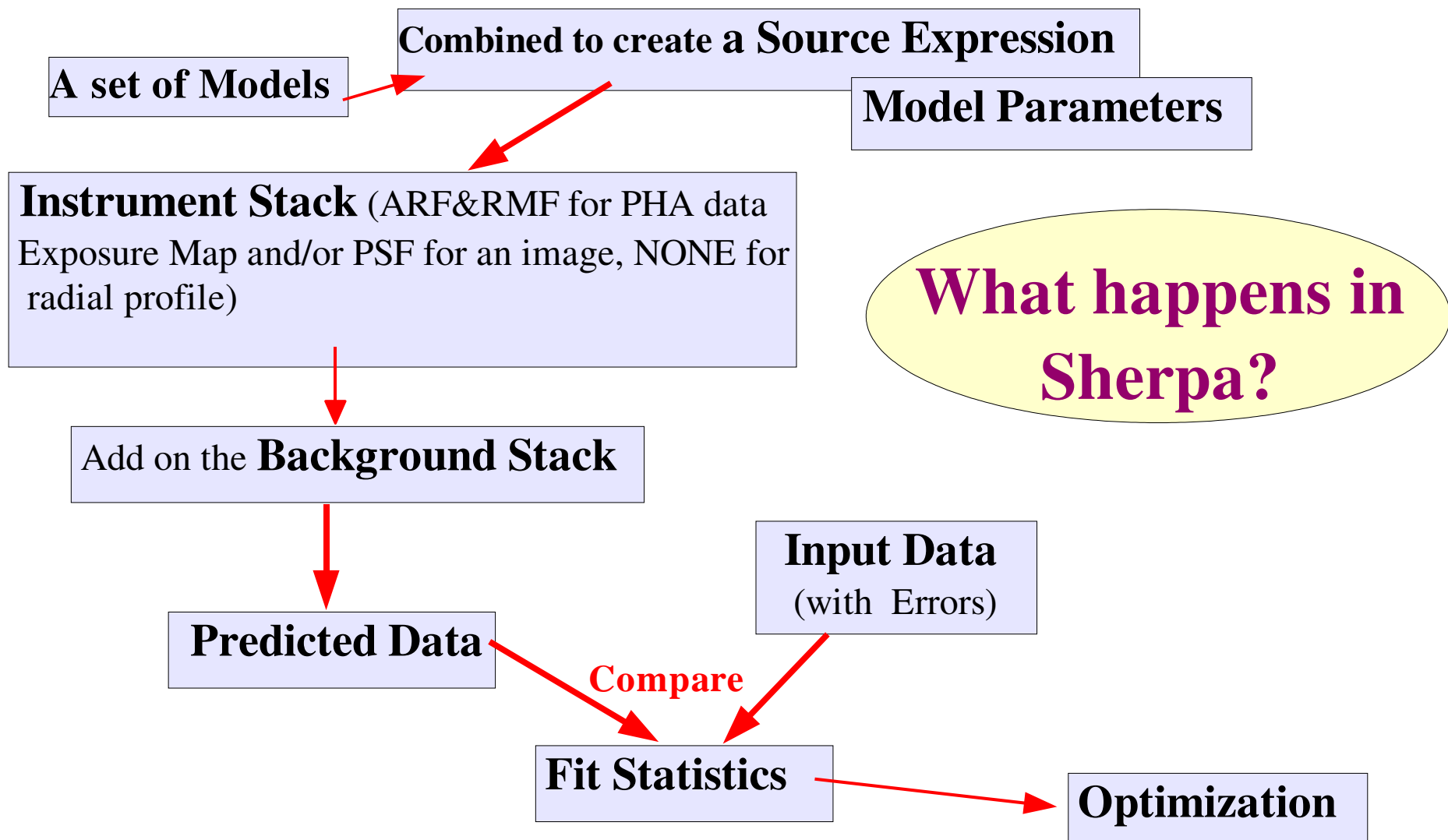
What can you do in Sherpa?

- Standard PHA based analysis.
- Model data in many spectral bands simultaneously, e.g., optical/ X-rays.
- Access ATOMDB and GUIDE/ISIS for grating data analysis.
- Fit radial profiles.
- Simulate 1D data.
- Model 2D image data, e.g., fit surface brightness of the extended source.
- Get normalization of your PSF, while fitting the data with 1D/2D PSF.
- Use the PSF as a convolution kernel in the 2D image analysis(FFT or sliding cell).
- Convolution using the TCD library kernel.
- Use of exposure maps in the image analysis.
- Joint-mode data: spatial-spectral, spatial-timing
- Use scripts based on Sherpa only commands.
- Use S-lang on command line and in S-lang based scripts.
- Use your own models with User Models and S-lang user models.



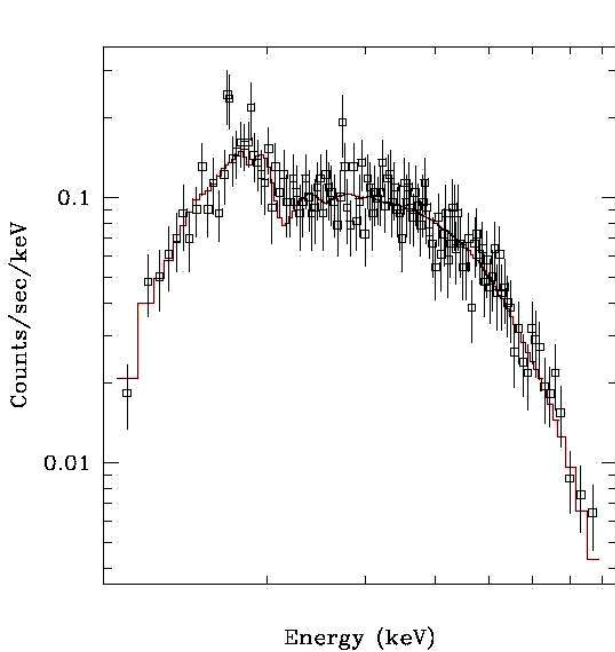
Standard PHA based analysis:

- Source data:
 - can be modeled in energy/wavelength space.
 - multiple data sets can be modeled with the same or different models in one Sherpa session.
 - data can be **filtered** on the command line, or from **filter** file.
- Instrument responses (RMF/ARF):
 - are entered independently from the source data.
 - one set of instrument responses can be read once and applied to multiple data sets.
 - several instrument responses used in analysis of one source model or multiple data sets.
 - multiple response files can be used in one source model expression.
- Background files:
 - are entered independently from the source data.
 - multiple background files can be used for one data set, e.g. grating analysis
 - the same background can be applied to multiple data sets.
 - background can be modeled independently of the source data, and have its separate instrument responses.
 - background can be modeled simultaneously with the source data.
 - background can be subtracted from the source data (subtract/unsubtract).

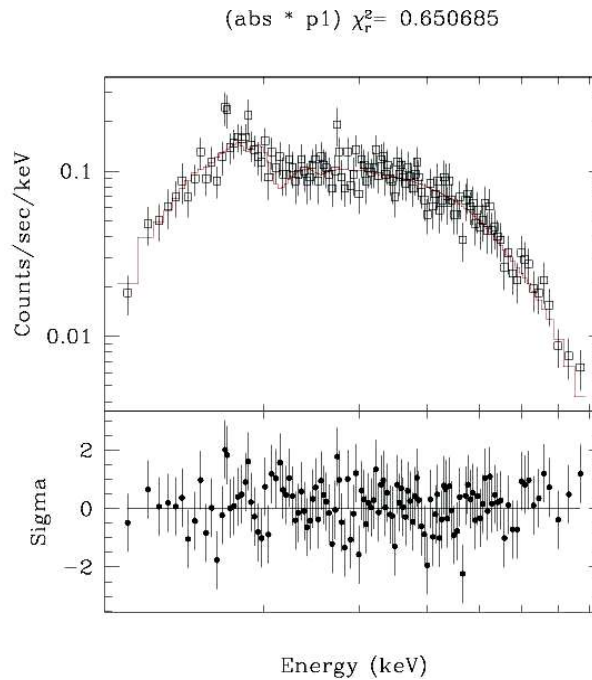




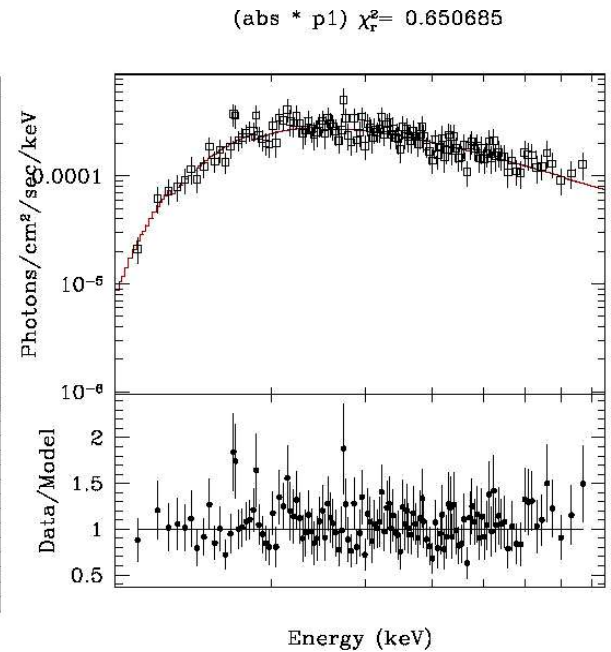
Displaying the Results



lp fit



lp 2 fit delchi



lp 2 ufit ratio



Main SHERPA Components

- Data Input/Output.
- Visualization through ChIPS and ds9
- Model library and model language.
- Statistics and Error Analysis.
- Optimization Methods.



Data Input/Output

- General use of data type and dimensionality.
- Supported types of files: ASCII, FITS binary tables and Images, PHA types I & II, IRAF IMH and QPOE files
- Sherpa:
 - ◆ groups the data if appropriate;
 - ◆ treats integer, float or double precision data;
 - ◆ supports data of arbitrary dimensionality
- I/O interface through Data Model and Varmm
- Filtering while reading the data.
- Input data on the command line in two ways.



```
s her pa> dat a "i mage. f i t s ( 1 5 0 : 3 0 0 , 1 6 0 : 3 1 0 ) "  
s her pa> s h o w  
C u r r e n t   D a t a   F i l e s :  
D a t a   1 :   i m a g e . f i t s ( 1 5 0 : 3 0 0 , 1 6 0 : 3 1 0 )   f i t s .  
T o t a l   S i z e :   2 2 8 0 1   b i n s   ( o r   p i x e l s )  
D i m e n s i o n s :   2  
S i z e :   1 5 1   x   1 5 1  
T o t a l   c o u n t s   ( o r   v a l u e s ) :   2 0 7 1 1   c t s
```

or

```
s her pa> mydat a=r ead f i l e (" i m a g e . f i t s ( 1 5 0 : 3 0 0 , 1 6 0 : 3 1 0 ) ")  
s her pa> pr i n t ( mydat a)  
_ f i l e n a m e           =   i m a g e . f i t s  
_ p a t h                 =   / d a t a /  
_ f i l t e r             =   ( 1 5 0 : 3 0 0 , 1 6 0 : 3 1 0 )  
_ f i l e t y p e         =   1 1  
_ h e a d e r             =   S t r i n g _ T y p e ( 2 6 8 )  
_ t r a n s f o r m       =   T A N  
_ n a x e s               =   2  
p i x e l s              =   F l o a t _ T y p e ( 1 5 1 , 1 5 1 )  
m i n                   =   D o u b l e _ T y p e ( 2 )  
m a x                   =   D o u b l e _ T y p e ( 2 )  
s t e p                  =   D o u b l e _ T y p e ( 2 )  
c r v a l                =   D o u b l e _ T y p e ( 2 )  
c r p i x                =   D o u b l e _ T y p e ( 2 )  
c r d e l t              =   D o u b l e _ T y p e ( 2 )  
s her pa> pr i n t ( mydat a . c r v a l ( 0 ) )  
2 7 8 . 3 8 6  
s her pa> pr i n t ( mydat a . c r v a l ( 1 ) )
```



MODELS

- Three main type of models:
 - ◆ **Source**
 - ◆ **Background**
 - ◆ **Instrument**
- Model library consists of several models (plus XSPEC v.11) which can be used to define a **source** or **background** model
- There are different types of **instrument models** to support both 1D and 2D analysis.
- **Instrument** models are **convolved** with **Source** and **Background** models before the model predicted data is compared with the observed data.
- Instrument and Background models are **NOT** required. Source models **have to be defined** for fitting.



Instrument Models

FARF1D FARF	1-D file-based ARF
FARF2D FEXPMAP FEXPMAP2D	2-D file-based Exposure Map
FPSF1D	1-D file-based PSF
FPSF2D FPSF PSFFROMFILE	2-D file-based PSF
FRMF1D FRMF	1-D file-based RMF
RSP1D RSP	1-D instrument model (ARF/RMF)
RSP2D	2-D instrument model (ExpMap/PSF)
TPSF1D	1-D TCD-model-based PSF
TPSF2D TPSF PSFFROMTCD	2-D TCD-model-based PSF



r sp1d[myfile]

	Param	Type	Value
1	rmf	string:	"file.rmfile"
2	arf	string:	"file.arfile"

tpsf2d[tpsf2d]

	Param	Type	Value	Min	Max
1	xsize	frozen	4	1	1024
2	ysize	frozen	4	1	1024
3	nsigma	frozen	2	1e-02	100
4	funcTyp	frozen	1	0	7
5	fft	frozen	1	0	1

The Function Type is: Gaussian.

fpsf2d[psf0]

	Param	Type	Value	Min	Max
1	file	string:	"psf_image.fits"		
2	xsize	frozen	32	1	1024
3	ysize	frozen	32	1	1024
4	xoff	frozen	0	-512	512
5	yoff	frozen	0	-512	512
6	fft	frozen	1	0	1



Instrument Models Expressions

```
sherpa> farf1d[a](arf.fits)
sherpa> frmf1d[r](rmf.fits)
sherpa> instrument = a*r
```

Here, the photon spectrum y is multiplied by the ARF, then folded through the RMF. This instrument stack is equivalent to

```
sherpa> instrument = rsp[a](rmf.fits,arf.fits)
```

Sets of instrument models separated by the + operator each fold the same evaluated photon spectrum y , with the resulting group of counts spectra being summed.

```
sherpa> farf1d[a1](arf_order1.fits)
sherpa> farf1d[a2](arf_order2.fits)
sherpa> frmf1d[r1](rmf_order1.fits)
sherpa> frmf1d[r2](rmf_order2.fits)
sherpa> instrument = a1*r1 + a2*r2
```



Model Language

- All predefined in model library models can be used in model expression to build a **source or background model**
- Each library model can be given a **unique name** within Sherpa session.

```
sherpa> gauss1d(g1)
sherpa> source = ATTEN(att1) * BPL(b1)
att1.hcol parameter value (1e+20)
att1.height ratio parameter value (0.1)
att1.height ratio parameter value (0.01)
b1.gamma1 parameter value (0)
b1.gamma2 parameter value (0)
b1.eb parameter value (100)
b1.ref parameter value (1)
b1.ampl parameter value (1)
```

- Model Parameters can be **linked** to other model parameters, arithmetic expression or other models.

```
sherpa> source = POLY(con) + gauss1d(g1) + gauss1d(g2)
sherpa> g1.ampl => 0.4 * g2.ampl
```

or

```
sherpa> func = const1d(red)
sherpa> g1.pos => 0.568 * func
```



? An argument of a model (e.g. energy) is defined as an expression in **Nested Models**.

Parameter Expression:

```
s her pa> Temper at ur e = PQLY
s her pa> BB. kT => Temper at ur e
s her pa> s how s our ce
BB
```

bbody(BB) (i n t e g r a t e : o n)		Par am	Type	Val ue	M i n	Max
		-----	-----	-----	---	---
1	kT	link	var yi ng	exp r e s s i o n : T e m p e r a t u r e		
2	ampl	t h a w e d	0. 3	1e- 20	3. 4028e+38	

Argument Expression:

```
s her pa> xener gy = SHL OG( mod)
s her pa> s our ce = BB{ xener gy}
s her pa> s how s our ce
```

BB{xenergy }

bbody(BB) (i n t e g r a t e : o n)		Par am	Type	Val ue	M i n	Max
		-----	-----	-----	---	---
1	kT	t h a w e d	0. 3	0. 1000	1000	
2	ampl	t h a w e d	0. 001	1e- 20	3. 4028e+38	

shl oge(mod) (i n t e g r a t e : o f f)

Par am	Type	Val ue	M i n	Max
-----	-----	-----	---	---
1	o f f s e t	f r o z e n	0- 3. 4028e+38	3. 4028e+38
2	c o e f f	f r o z e n	1- 3. 4028e+38	3. 4028e+38
3	a m p l	f r o z e n	1	0
				3. 4028e+38



- For **Joint-Mode** analysis one can apply models on each axis:

```

s her pa> DATA image.fits FITS IMAGE
s her pa> LORENTZ(Spatial Axis 0) (98: 5: 200, 70: 50: 90, 1: 1: 200)
s her pa> POWLAW(D(SpecAxis 1))
s her pa> SRC = Spatial Axis 0{x1} * SpecAxis 1{x2}
s her pa> show source
(Spatial Axis 0{ 0 } * SpecAxis 1{ 1 })

```

lorentz1d(Spatial Axis 0) (integration: on)

	Par am	Type	Val ue	Mi n	Max
	-----	-----	-----	---	---
1	whm	thawed	98	5	200
2	pos	thawed	70	50	90
3	ampl	thawed	1	1	200

powlaw1d(SpecAxis 1) (integration: on)

	Par am	Type	Val ue	Mi n	Max
	-----	-----	-----	---	---
1	gamma	thawed	1.5	-10	10
2	ref	frozen	1-3.4028e+38	3.4028e+38	
3	ampl	thawed	1	1e-20	3.4028e+38



Fit Statistics in Sherpa:

Sherpa has a large array of statistics appropriate for analyzing Poisson-distributed (*i.e.* counts) data.

- Statistics based on χ^2 :
 - CHI GEHRELS
 - CHI DVAR
 - CHI MVAR
 - CHI PARENT
 - CHI PRIMINI
- Statistics based on the Poisson likelihood:
 - CASH
 - BAYES

If the data are not Poisson-distributed (*i.e.* fluxes), then alternatives include:

- ? least-squares fitting: setting all variances to one
- ? providing errors in an input file.



Optimization in Sherpa

Optimization => minimizing the statistics (χ^2 or $\log \mathcal{L}$) by varying the thawed parameters of the model.

- **Find a local minimum:**

LEVENBERG-MARQUARDT
POWELL
SIMPLEX



Fast, but not appropriate for finding the global minimum of a complex statistical space when starting from a random point

- **Attempt to find the global minimum:**

GRID
GRID-POWELL
MONTECARLO
MONTE-LM
MONTE-POWELL
SMULATED ANNEALING



Computationally intensive algorithms designed to search complicated statistical surfaces.

- **Optimize/Reject/Filter:**

SIGMA-REJECTION outliers are filtered from the data.



Confidence Intervals

- Vary a parameter's value, while holding the values of all the parameters to their best-fit values, until the fit statistic increases by some preset amount from its minimum value ($\chi^2 = 1$ for 1).
 - Uncertainty
 - Projection
- Calculate **Covariance** matrix:

1 confidence intervals are given by $\sqrt{C_{i,i}}$

where $C_{j,i} = I_{i,j}^{-1}$

and $I_{i,j}$ - the information matrix computed at the best-fit point:

$$I_{i,j} = \frac{\partial^2 \chi^2}{\partial p_i \partial p_j}$$

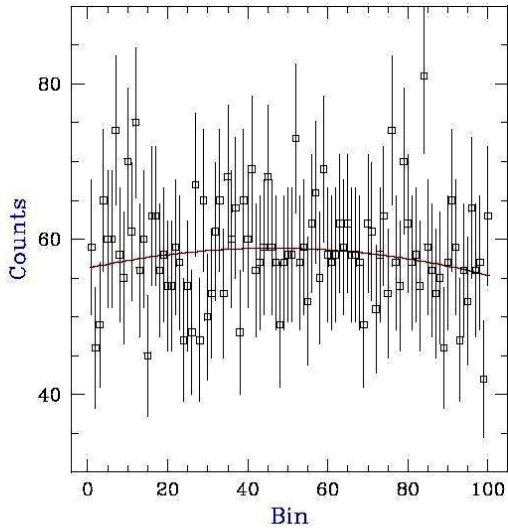
or any other statistics



Visualize Confidence Levels

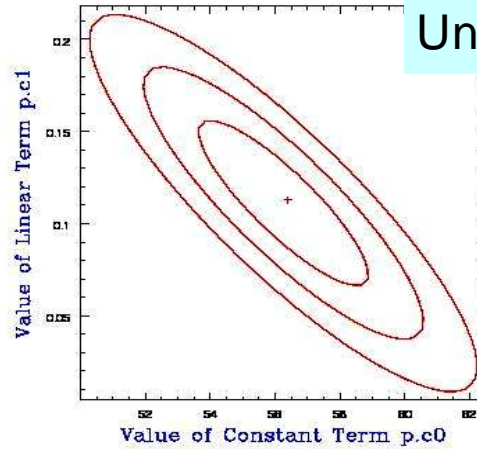
Uncertainty

Data and the Best Fit Model



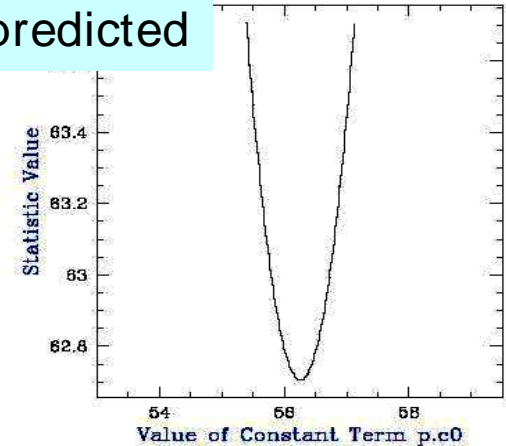
Well behave parameter space!

Confidence Region - Uncertainty

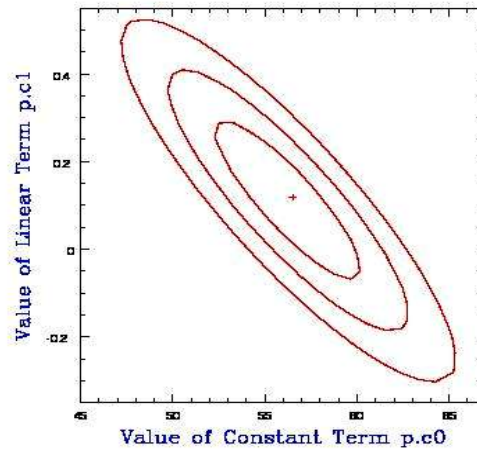


Underpredicted

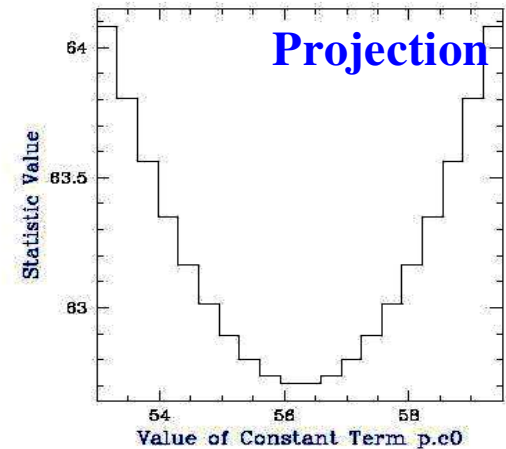
Interval - Uncertainty



Confidence Region - Projection



Interval - Projection



Projection



Customize Sherpa

- **Sherpa State Object** (e.g. Configuration file) – S-lang variable initialized at the start of the Sherpa session:

```
sherpa> print(sherpa)
plot          = sherpa_Plot_State
dataplot      = sherpa_Plot_State
fitplot       = sherpa_FitPlot_State
resplot       = sherpa_Plot_State
multiplot     = sherpa_Draw_State
output        = sherpa_Output_State
regproj       = sherpa_VisParEst_State
regunc        = sherpa_VisParEst_State
intproj       = sherpa_VisParEst_State
intunc        = sherpa_VisParEst_State
proj          = sherpa_Proj_State
cov           = sherpa_Cov_State
unc           = sherpa_Unc_State
con_levs      = NULL
modeloverride = 0
multiback     = 0
deleteframes  = 1
clobber       = 0
```

Customize Plotting

Customize Confidence Levels

```
sherpa> print(sherpa.regproj)
fast          = 1
expfac        = 3
arange        = 1
min           = Double_Type[2]
max           = Double_Type[2]
log           = Integer_Type[2]
nloop         = Integer_Type[2]
sigma         = Double_Type[3]
```



Customize Sherpa

- **Sherpa Resource File:**
 - a text file with Sherpa/Chips/S-lang commands
- **Access:**
 - Environment variable **SHERPARC**
 - File **.sherparc** in current directory \$PWD
 - File **.sherparc** in HOME directory \$HOME
- **Example:**

```
unix% more .sherparc
# Example Sherpa resource file
message("Starting to process sherparc")
paramprompt off
method simplex
define q () { () = sherpa_eval("quit"); }
message("Finished processing .sherparc")
```



```
unix% sherpa
...
  Abundances set to Anders & Grevesse
Starting to process .sherparc
Model parameter prompting is off
Finished processing .sherparc
sherpa> show method
Optimization Method: Simplex
  Name      Value      Min      Max      Description
  ----      -
1  iters     2000       1      1 0000  Maximum number of iterations
2  eps       1e-03     1e-04     100    Absolute accuracy
3  alpha      1          0.1       2    Algorithm convergence factor
4  beta       0.5       5e-02     1    Algorithm convergence factor
5  gamma      2          1.1      20    Algorithm convergence factor
sherpa> q
Goodbye.
```



Learn More on Sherpa Web Page



WHAT'S NEW | WATCH OUT

[Analysis Threads](#) | [Ahelp](#) | [Documents](#) | [Scripts](#) | [LEAO](#) | [ChART](#) | [CIAO](#)

Sherpa, CIAO's generalized modeling and fitting engine, allows users to construct complex models and to fit models to data in N dimensions. It has a library of optimization methods and fit statistics. *Sherpa* is "domain independent", i.e. it does not require particular axes to be fit. It is also mission independent, with no particular tie to Chandra data. For example, it has been used to analyze HST spectra.

Sherpa supports *S-Lang*, an interpreted programming language that can be used for scripting and data manipulation. Existing *S-Lang* scripts and utilities are available for download on the [CIAO scripts page](#).

The *GUIDE* package within *Sherpa* links *Sherpa* results (stored in a *MDL* file) to the *ATOMDB*, enabling the identification of spectral lines and the use of their properties in further fitting.

In order to run *Sherpa*, you must [download and install CIAO](#).

Sherpa CIAO 3.0 Highlights

- Multiple components in one instrument model expression are allowed. This supports, for example, the use of multiple response files in the analysis of LETG overlapping orders spectra.
- [Levenberg-Marquardt](#) (LM) is now *Sherpa*'s default optimizer.
- Two new optimization methods: [MONTE-LM](#) and [SIGMA-REJECTION](#)
- Configuration of the error estimation commands (e.g. [PROJECTION](#)) has moved from internal *Sherpa* functions to the *Sherpa* configuration variables (e.g., [sherpa.proj](#), [sherpa.regproj](#))
- *Sherpa* plot configuration is now done via the configuration variables (e.g. [sherpa.plot](#))
- Many new data access functions (e.g. [get_data](#), [load pha](#) etc.) have been added to the [Sherpa/S-Lang module](#), which has been enhanced considerably.
- ``import ("sherpa")`` allows for importing of the [Sherpa/S-Lang module](#) into other *S-Lang*-aware applications.

See the [Sherpa release notes](#) for a complete list of CIAO 3.0 changes.

<http://cxc.harvard.edu/sherpa/>

Sherpa Threads for CIAO 3.0

When running a thread for the first time, you may wish to follow along, using the actual data employed in the thread. Please see the [Getting Started](#) thread for instructions on how to download and use the example data.

All threads

A list of all the threads on one page.

Introductory

These threads cover the basics of *Sherpa*: reading data, establishing models, fitting, plotting, and basic customizations.

Fitting

Sherpa provides extensive facilities for modeling and fitting data. The topics here range from basic fits using source spectra and responses to more advanced areas such as simultaneous fits to multiple datasets, accounting for the effects of pileup, and fitting spatial and grating data.

Plotting

Sherpa allows the user to plot data, fits, statistics, ARFs, contours, surfaces, and more. These threads describe the basics of plotting as well as various methods for customizing plots.

Statistics

Sherpa provides numerous tools for determining goodness of fit, errors in parameter values, confidence intervals, and other statistical measures of a model's validity. These threads describe how to use these tools in your analysis.

S-Lang

The *S-Lang* language and [Sherpa/S-Lang module](#) provide a powerful means of extending *Sherpa*'s capabilities through custom-made functions and scripts. The threads here introduce *Sherpa*'s *S-Lang* functionality and provide some examples of its use.

Miscellaneous

These threads describe other tasks that one can perform using *Sherpa*.

Datasets

Links to the datasets used in the threads.