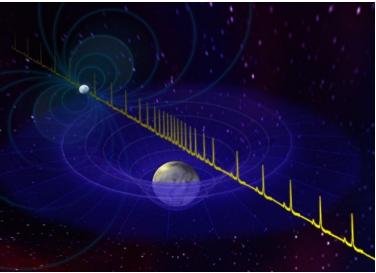


# Time evolution of luminosity and spin period of X-ray pulsars in the Small Magellanic Cloud

Jun Yang<sup>1,2</sup>

Silas Laycock<sup>1</sup>, Dimitris Christodoulou<sup>1</sup>, Jeremy Drake<sup>2</sup>, Samuel Fingerman<sup>1</sup>, Jaesub Hong<sup>2</sup>, Andreas Zezas<sup>2</sup>, Vallia Antoniou<sup>2</sup>, Malcolm Coe<sup>3</sup>, Wynn Ho<sup>3</sup>

1. University of Massachusetts, Lowell, MA, United States.
2. Harvard-Smithsonian Center for Astrophysics, Cambridge, MA, United States.
3. University of Southampton, Southampton, United Kingdom



# Outline

## ❖ Introduction to the library of pulsars in the Magellanic Clouds

event lists, light curves, periodograms & spectra

- Why study pulsars in Magellanic Clouds:

NSs hold unsolved problems; population-statistical study.

low absorption, minimum foreground contamination, known distance (->categorize luminosity of the Pulsars). The process of accretion and X-ray emission in HMXB

## ❖ Known pulsars with combination of 3 satellites, physical parameters

e.g. SXP 348, SXP 1323

## ❖ Pulse profile modeling

e.g. SXP 504

# Library of pulsars in the Magellanic Clouds

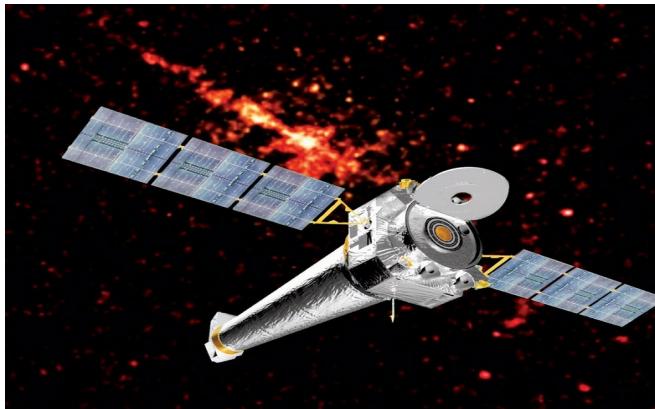
Pulsars In SMC & Model

Intro Library

Pulsar spin up /down

Pulse profile model

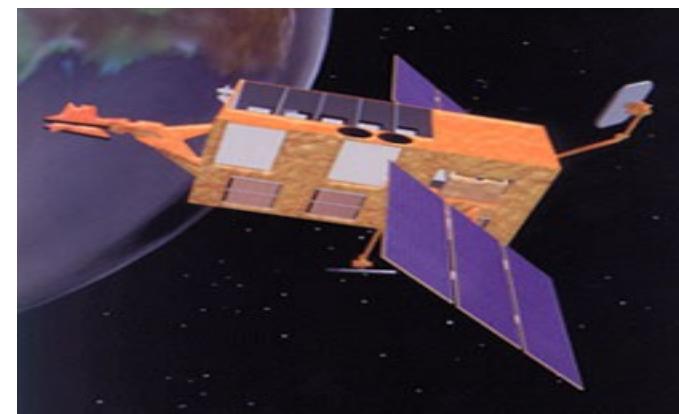
In sum



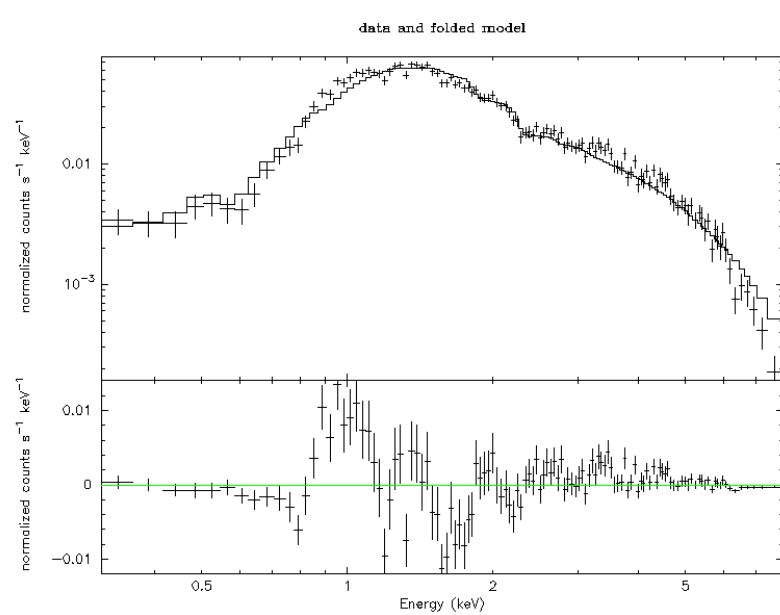
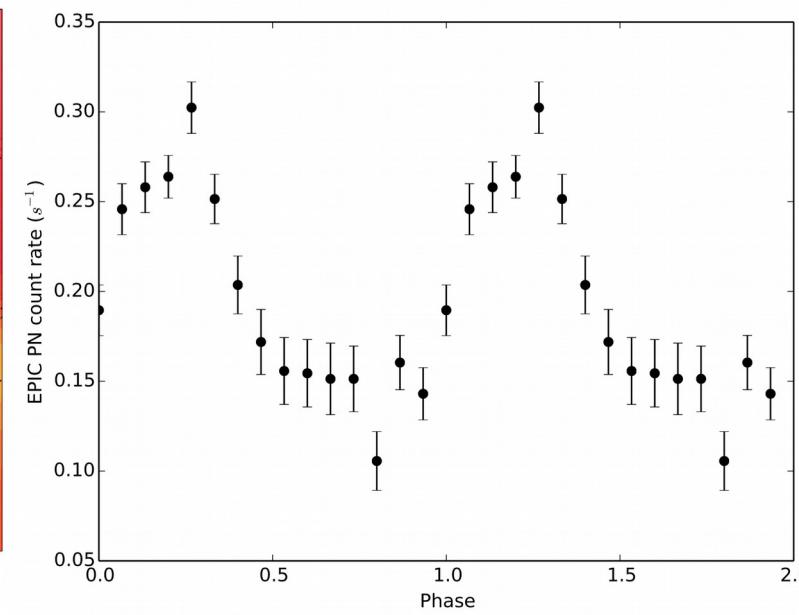
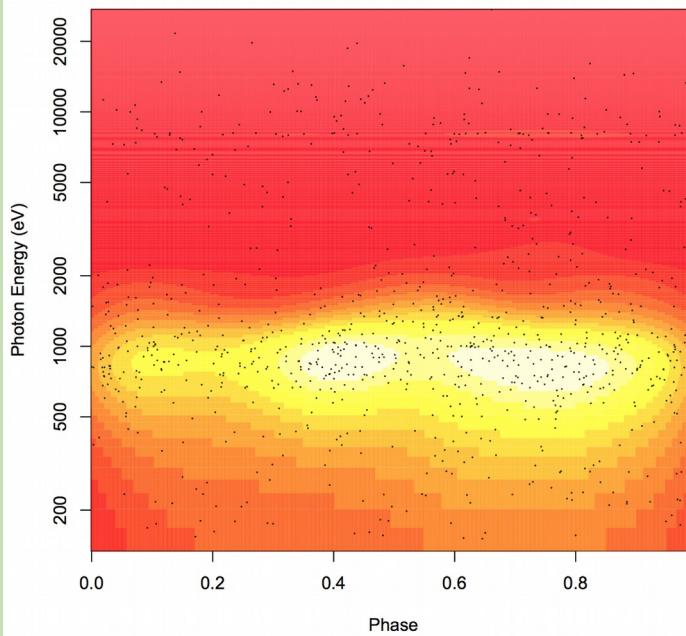
Chandra  
155+30 Obs: 2000-2014



XMM-Newton  
116+42 Obs: 2000-2014



RXTE  
Weekly Obs: 1997-2012



# Luminosity of X-ray pulsars in the Small Magellanic Cloud

Pulsars In SMC & Model

**Intro Library**

Pulsars spin up/down

Pulse profile model

In sum

Observation overview

Luminosity Range:

$10^{31}$ - $10^{38}$  erg/s

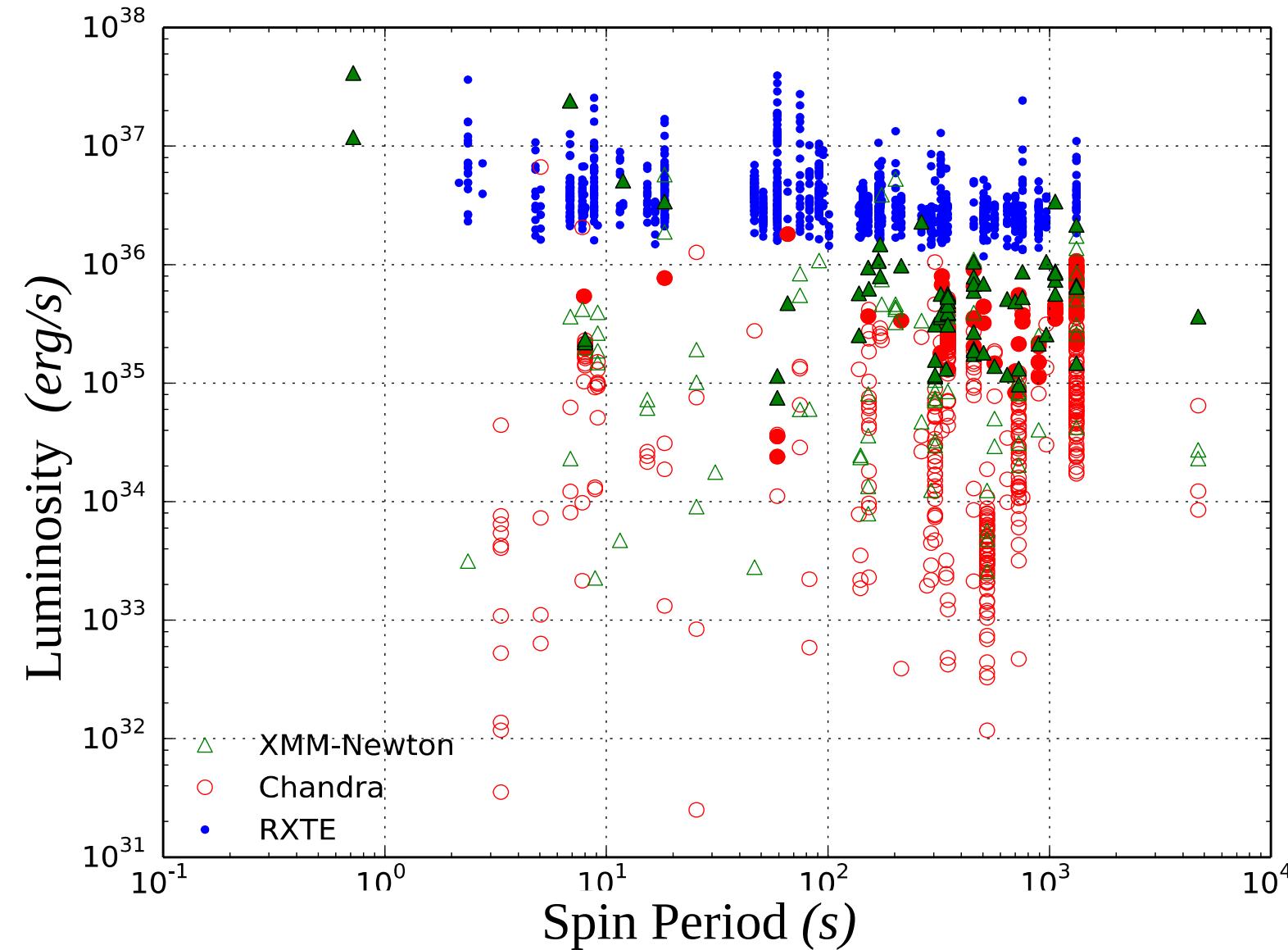
Fast pulsars

rarely detected

Faint source

Short exposure

Classify:  
pulsating & non-pulsating



# Luminosity of X-ray pulsars in the Small Magellanic Cloud

Pulsars In SMC & Model

**Intro Library**

[Pulsars spin up/down](#)

Pulse profile model

In sum

Observation overview

Luminosity Range:

$10^{31}$ - $10^{38}$  erg/s

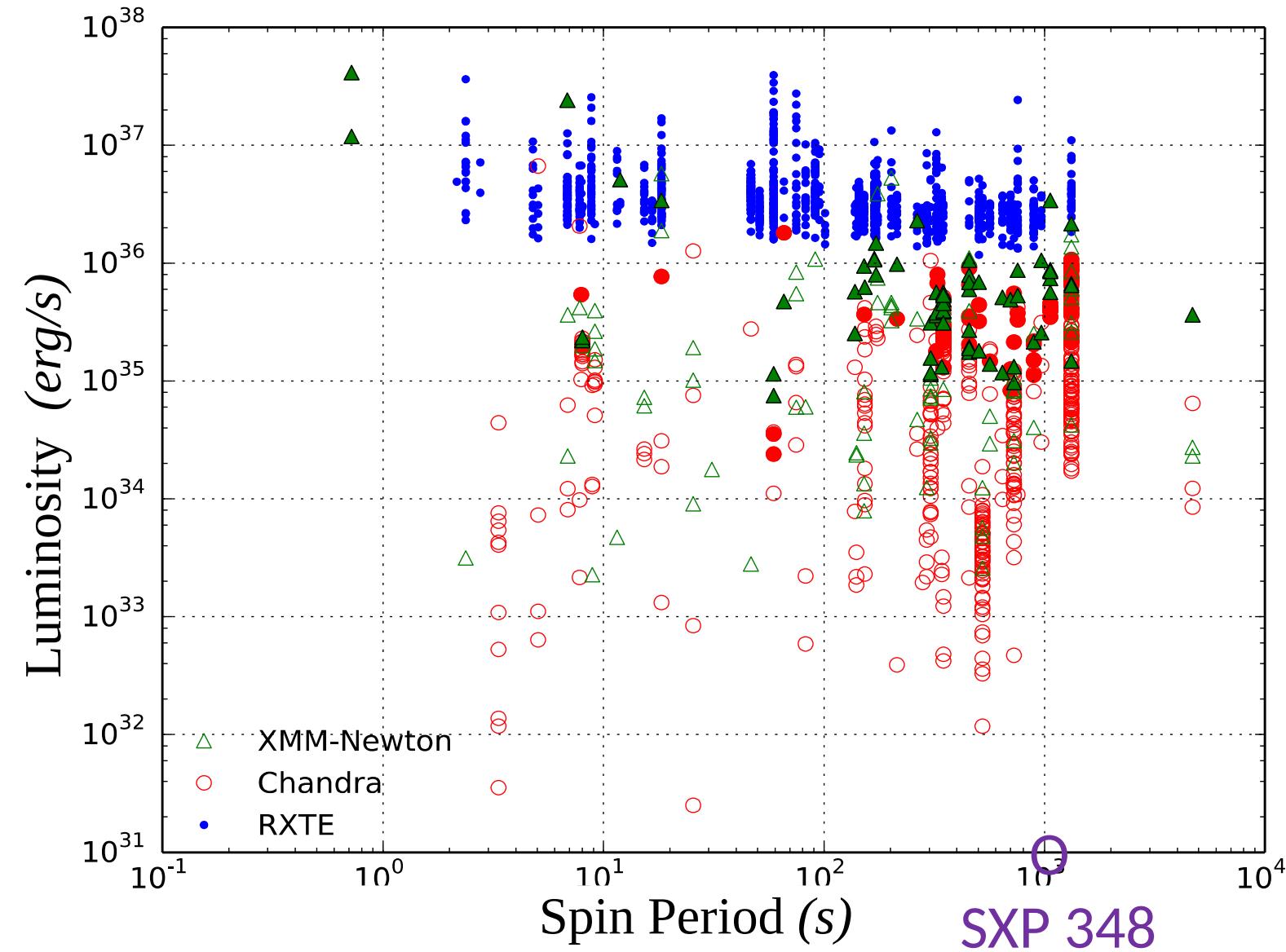
Fast pulsars

rarely detected

Faint source

Short exposure

Classify:  
pulsating &  
non-pulsati



# Time evolution of Source properties with 3 satellite combination

Pulsars  
In SMC  
&Mode  
I

Intro  
Library

Pulsa  
rs  
spin  
up  
/down

Pulse  
profile  
model

In sum

e.g.

SXP 348

Luminosi  
ty  
(erg/s)

*RXTE does  
not have  
Pulse Frac  
info*

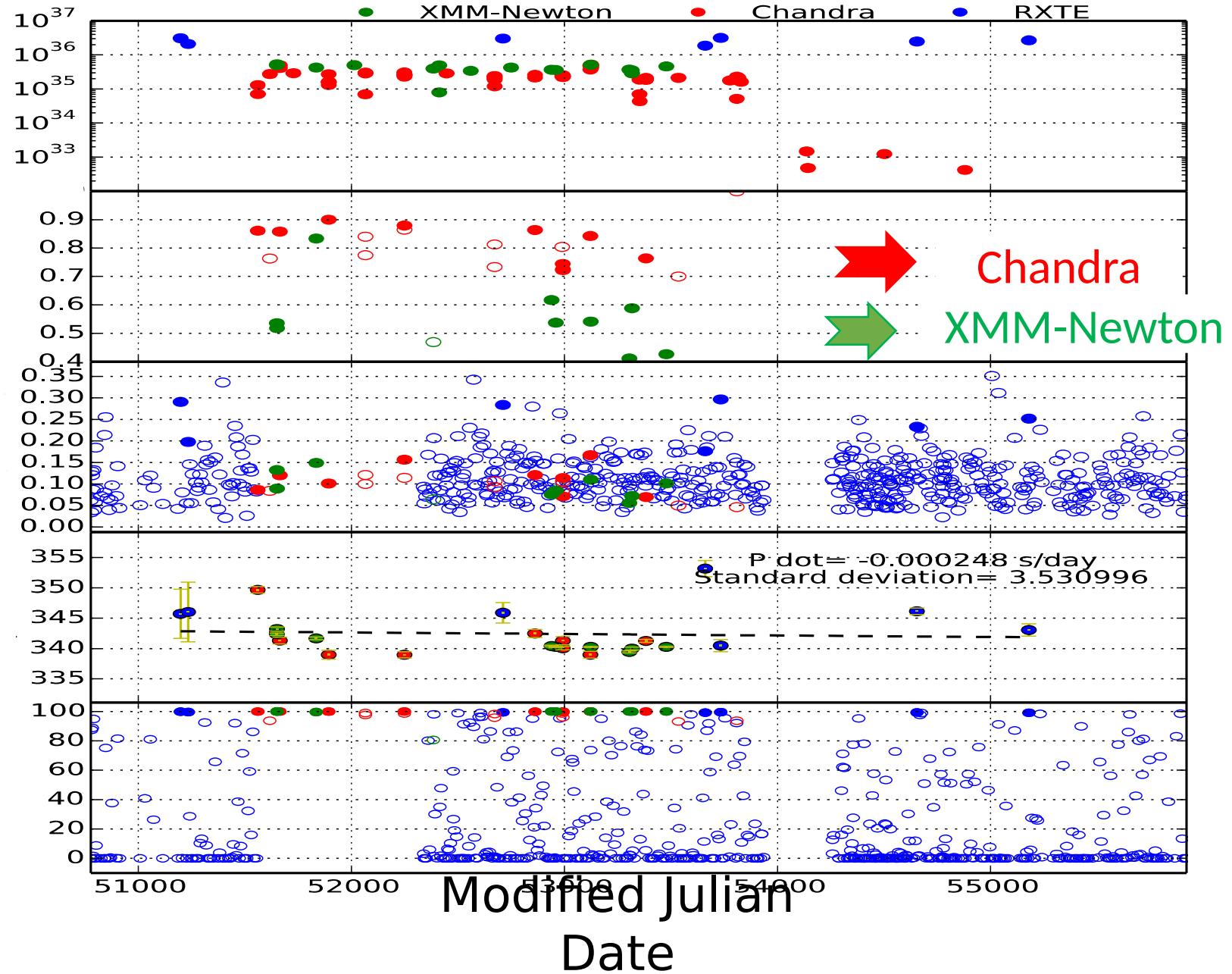
Pulse  
frac  
tion

Amplitud  
e  
(Counts/  
)

Spin  
Period(s  
)

*Spin up:*  
Mass and  
angular  
momentum to  
NS

Significan  
ce  
(%)



# SMC Pulsar library products

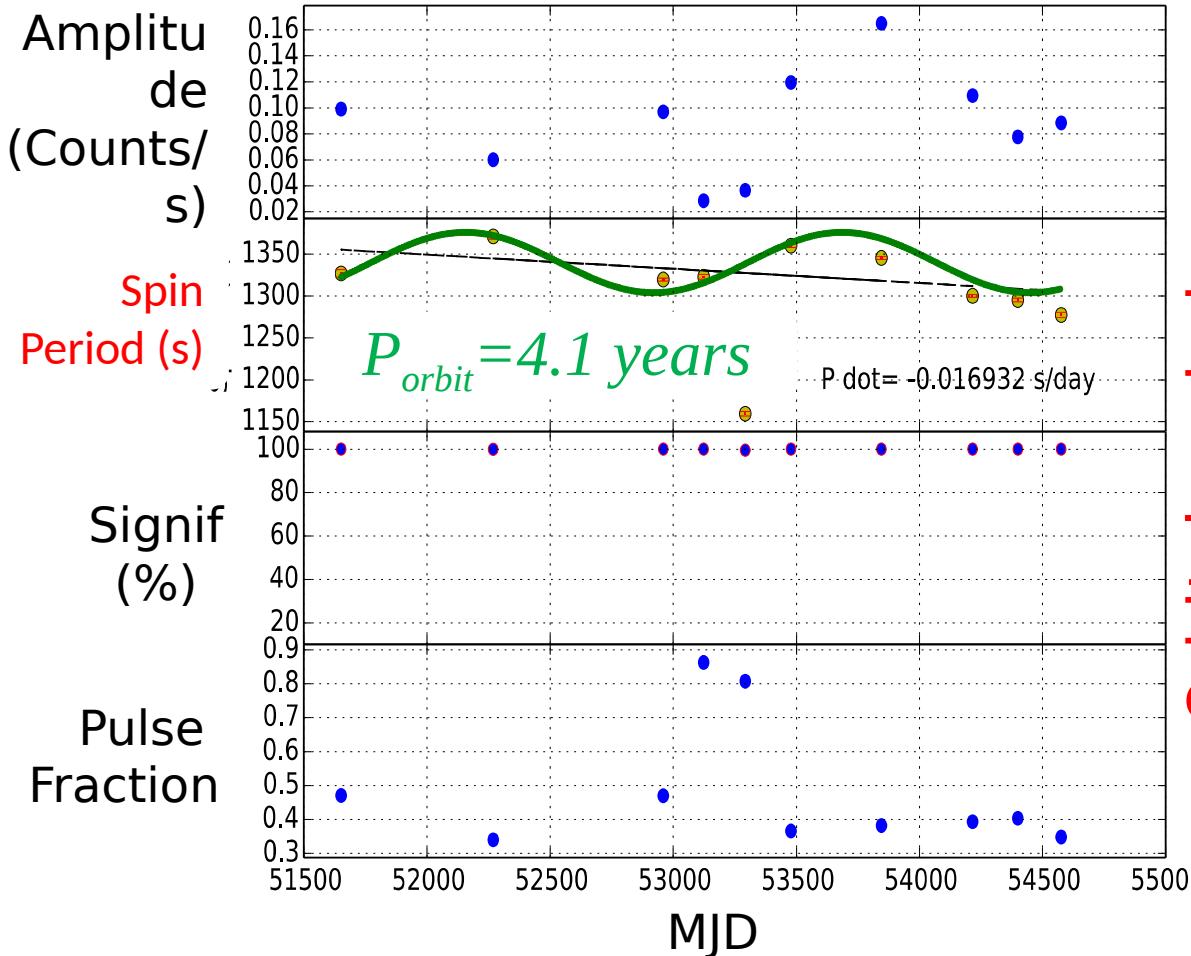
Pulsars  
In SMC  
&Mode  
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Intro  
Library

**Pulsa  
rs  
spin  
up  
/down**

Pulse  
profile  
model

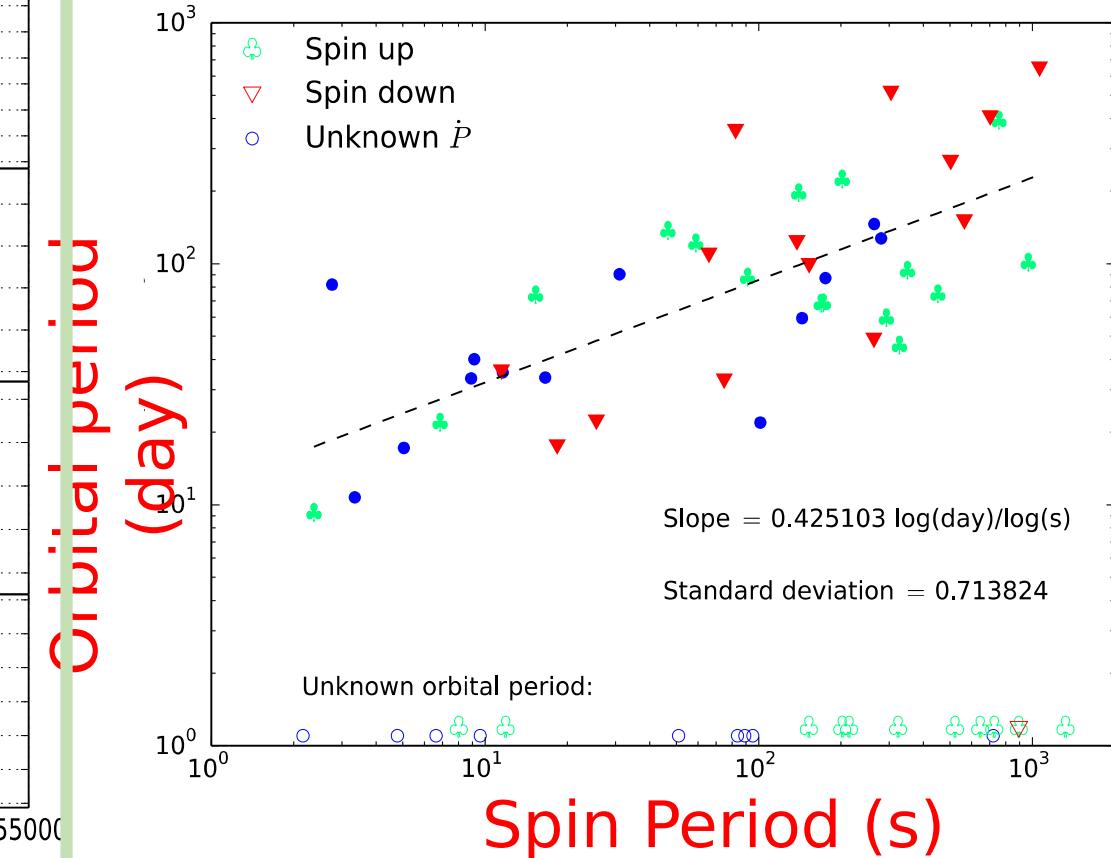
In sum



e.g. SXP 1323

Spin up: 0.016932 s/day

Orbital modulation, accretion torque



$$P_{\text{orbit}} = 12.0656 \text{ (days)} * P_{\text{spin}}^{0.4251}$$

26 spin up 17 spin down

43 out of 65 pulsars have measured orbit period

# Model: Off-center magnetic axis with GR effect

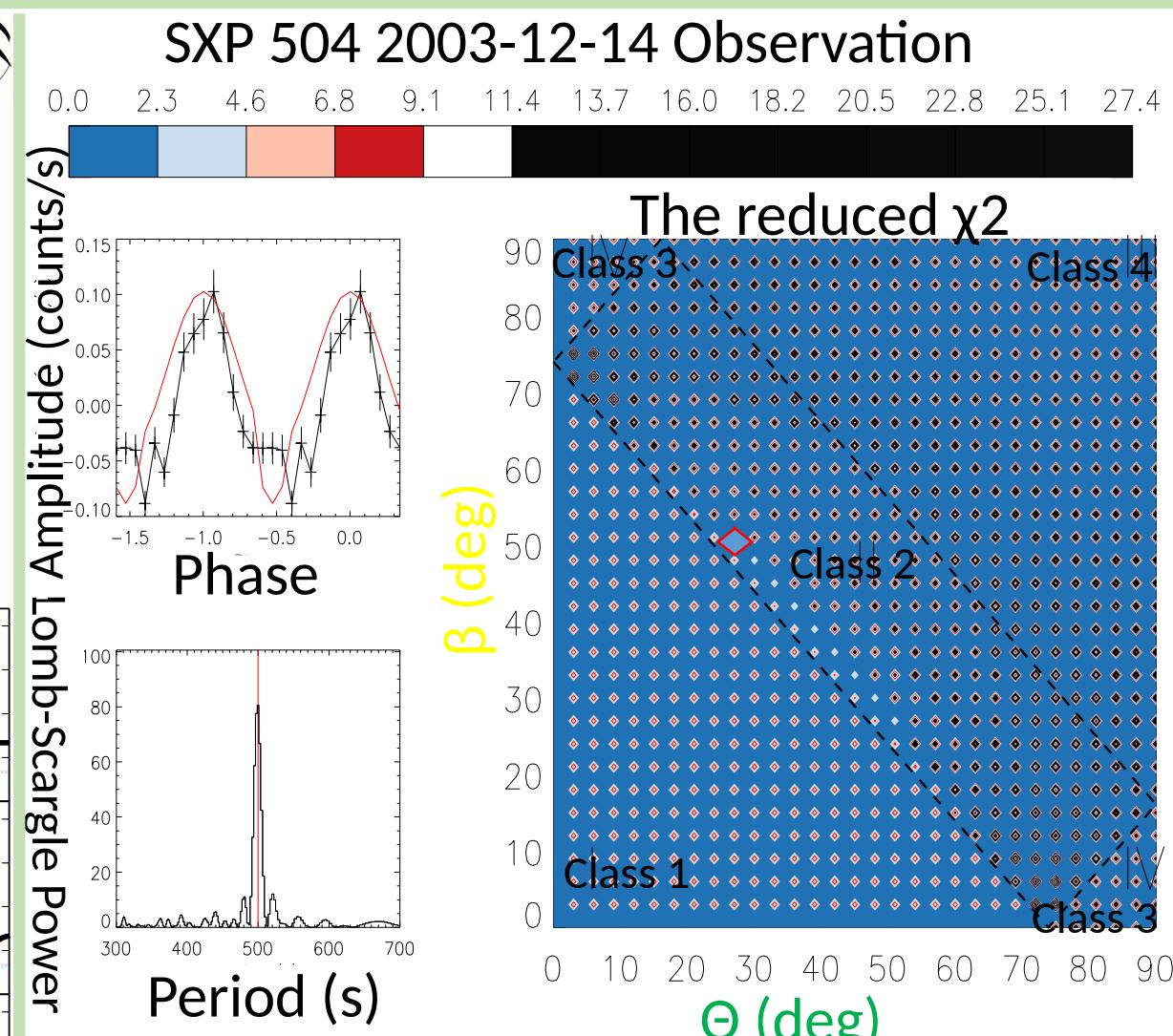
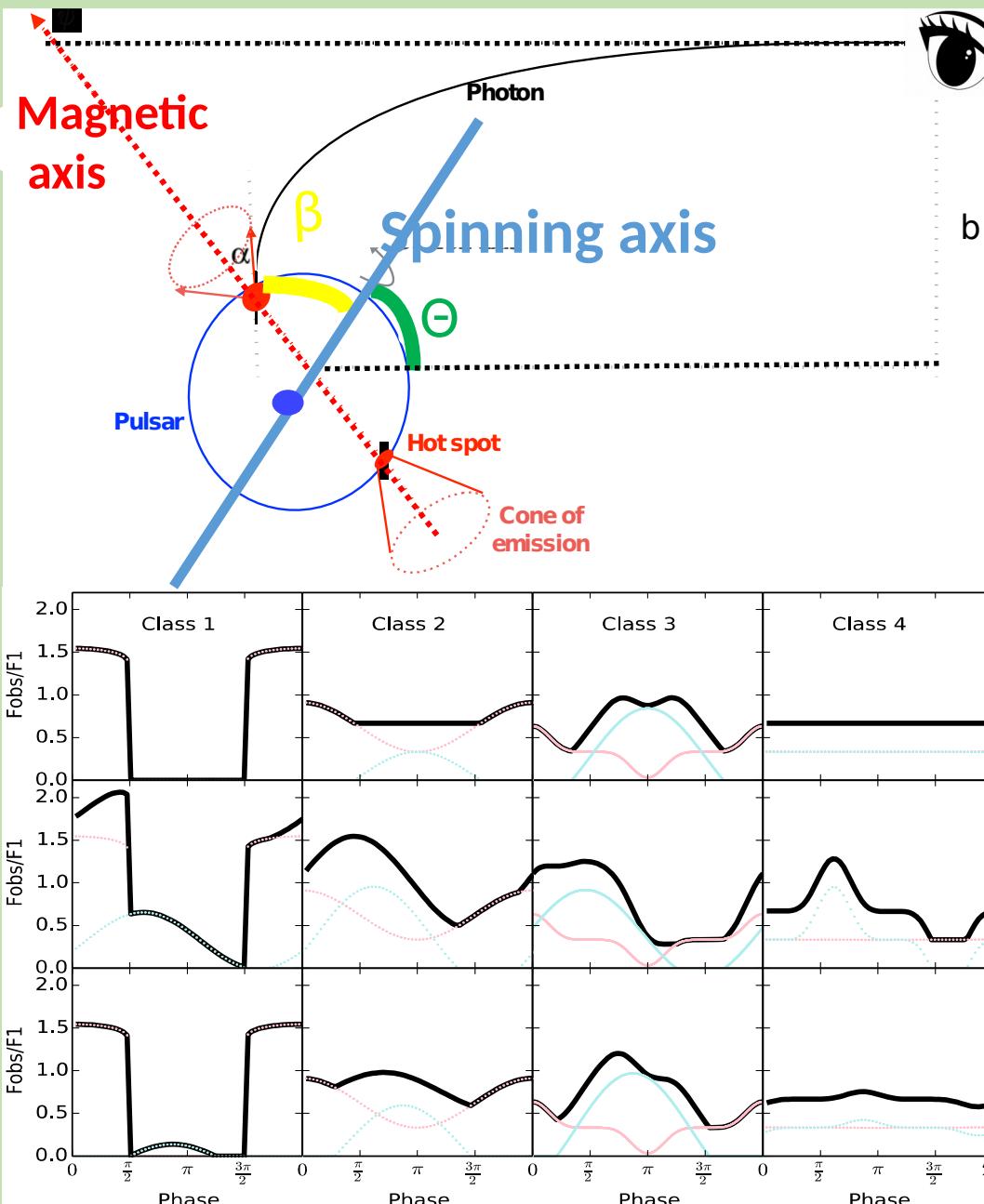
Pulsars  
In SMC  
&Mode  
I

Intro  
Library

Pulsar  
S  
spin up  
/down

Pulse  
profil  
e  
model

In sum



Best fitting angle:  $\Theta = 28^\circ$  and  $\beta = 52^\circ$

$\Theta$ : angle between spinning axis and B

$\beta$ : angle between spinning axis and light of sight

# Summary & Outlook

Pulsars  
In SMC  
&Mode  
I

- ❖ Library of 3 satellites combination: Chandra, XMM-Newton, RXTE  
Known pulsars in SMC & LMC : event lists, pulse profiles, spectra

Into  
Library

- ❖ Time evolution of their luminosities and spin period  
26 spin up and 17 spin down

Pulsar  
s  
spin  
up  
/down

- ❖ Modeling  
More phenomenon into current model: e.g., flow column geometry including physics of accretion shock, accretion rate and photon energy

Pulse  
profile  
model

- ❖ Astrosat & HXMT (Hard X-ray Modulation Telescope)

Investigate the hard X-ray sources; bright galactic pulsars, expand the period range of the pulsars

In  
sum

Pulsars In SMC & Model

Into Library

Pulsars spin up /down

Pulse profile model

In sum





# Library of pulsars in the Magellanic Clouds

Pulsars  
In SMC  
&Mode  
I

Intro  
Librar  
y

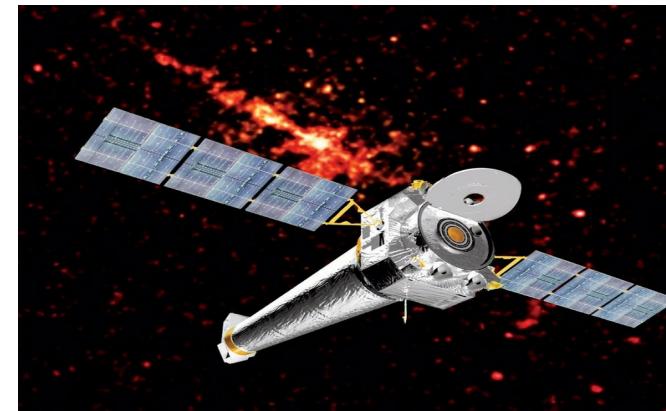
Pulsar  
ss  
pin up  
/down

Pulse  
profile  
model

In sum



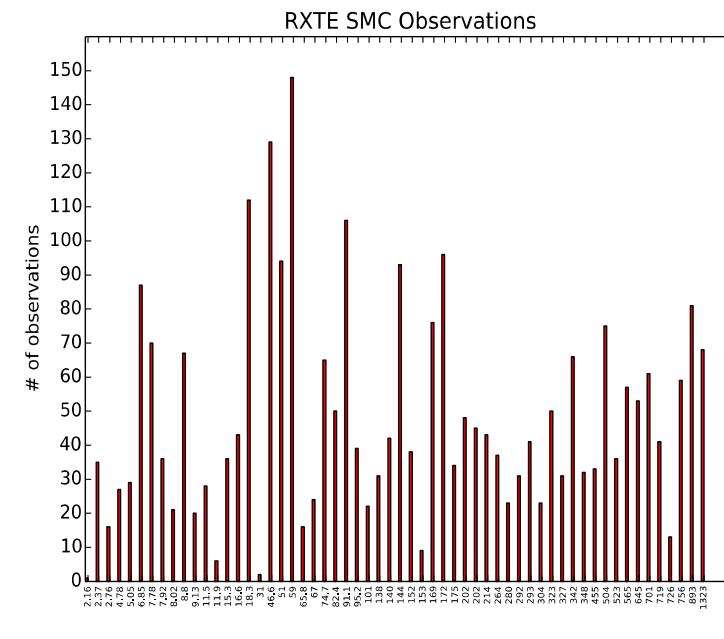
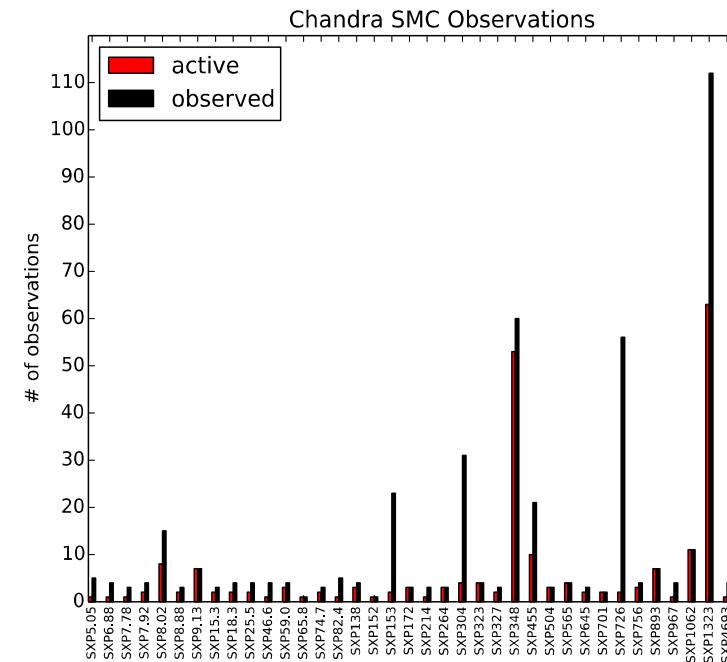
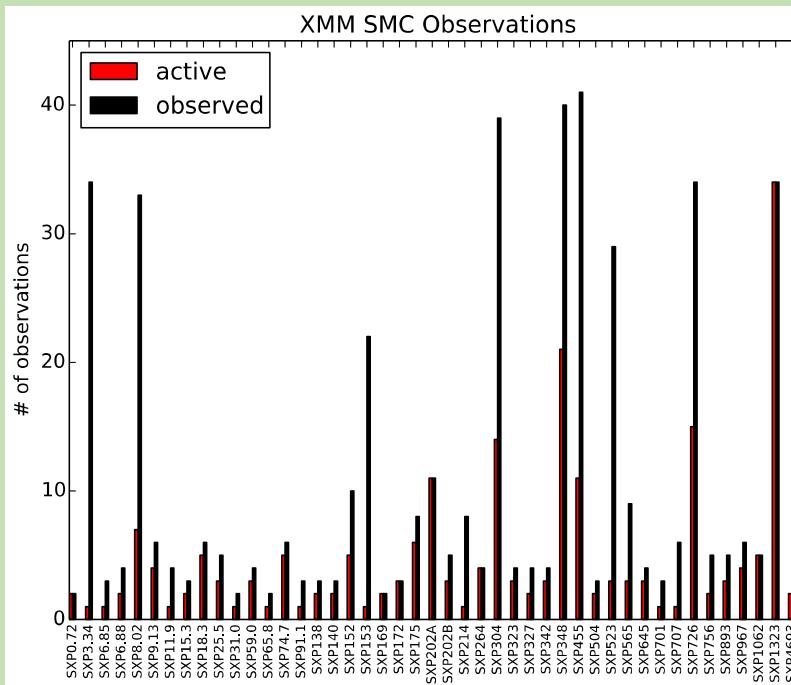
XMM-Newton  
116+42 Obs: 2000-2014



Chandra  
155+30 Obs: 2000-2014

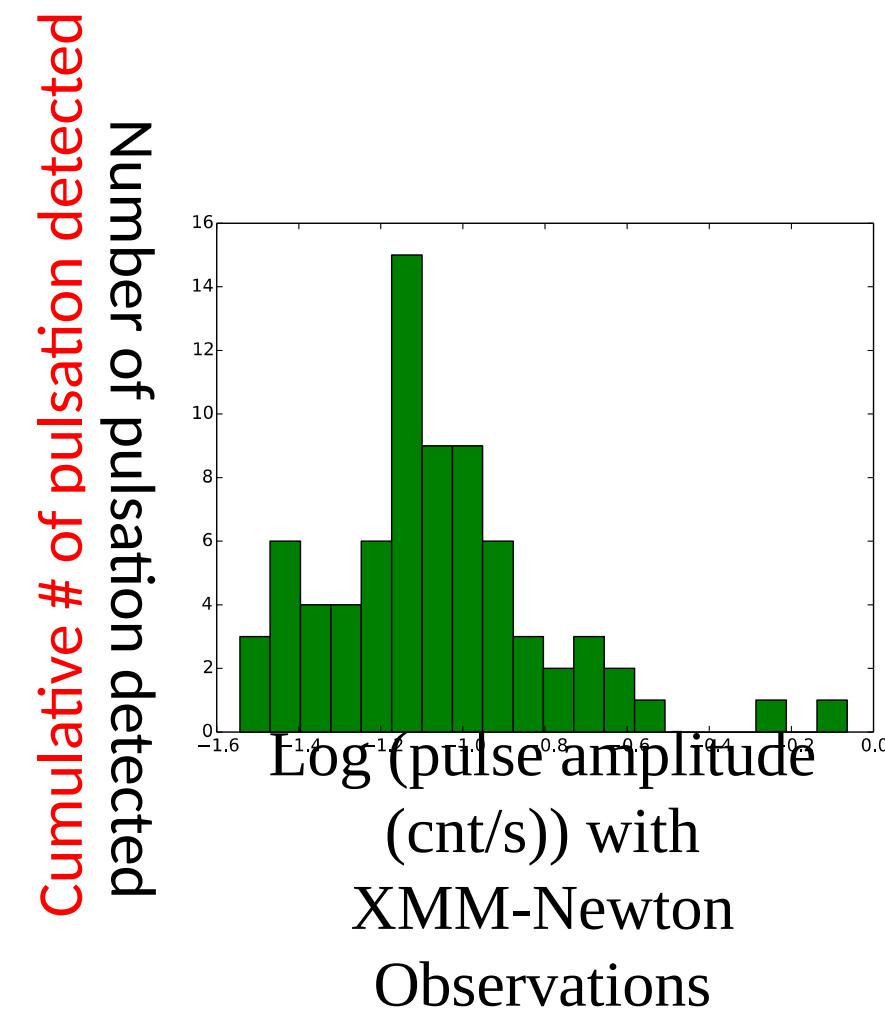
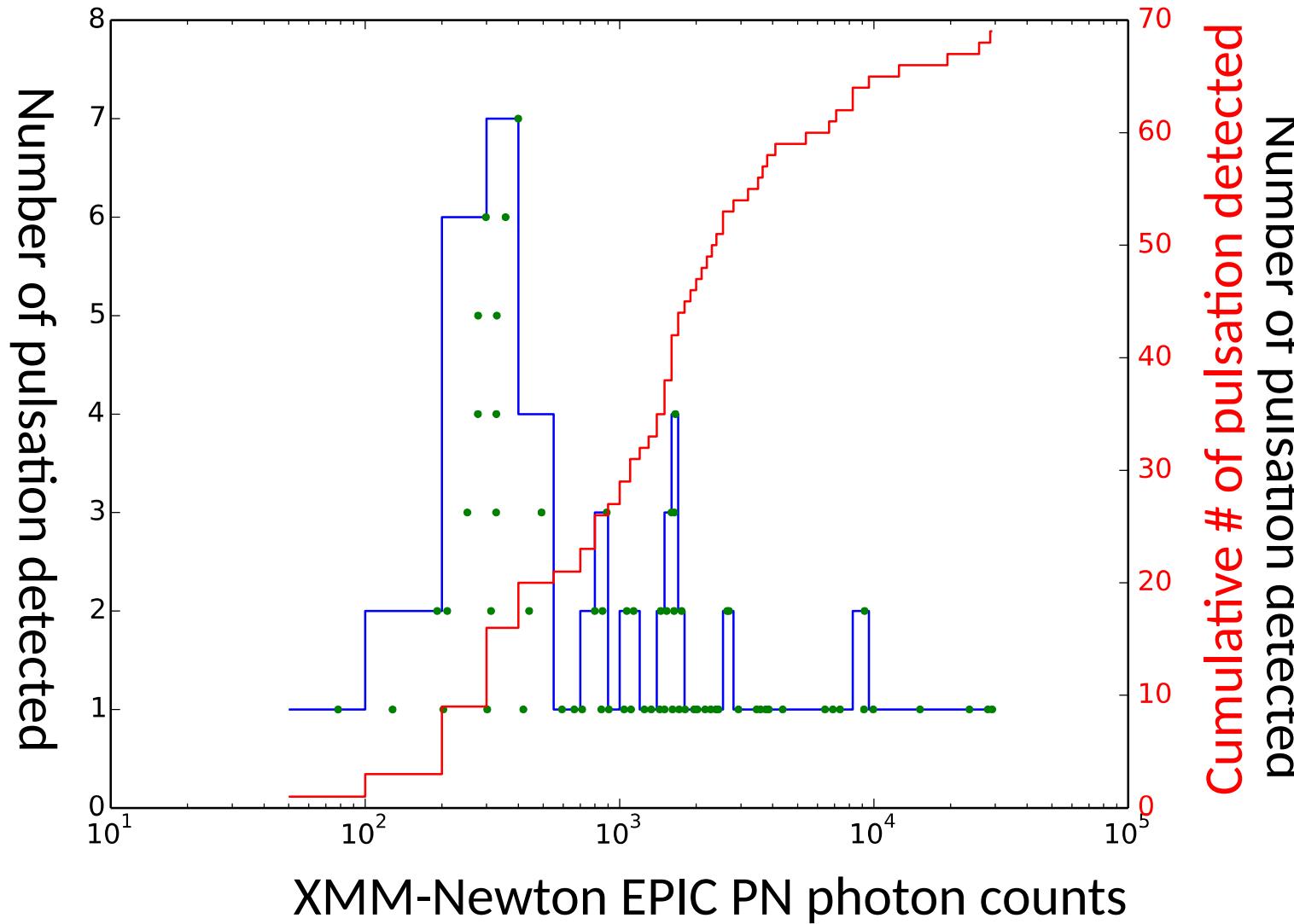


RXTE  
Weekly Obs: 1997-2008



# Library of pulsars in the Magellanic Clouds

Pulsars  
In SMC  
& Mode  
I  
  
Into  
Library  
  
Pulsar  
s  
spin  
up  
/down  
  
Pulse  
profile  
model  
  
In sum



# Library of pulsars in the Magellanic Clouds

Pulsars  
In SMC  
&Mode  
I

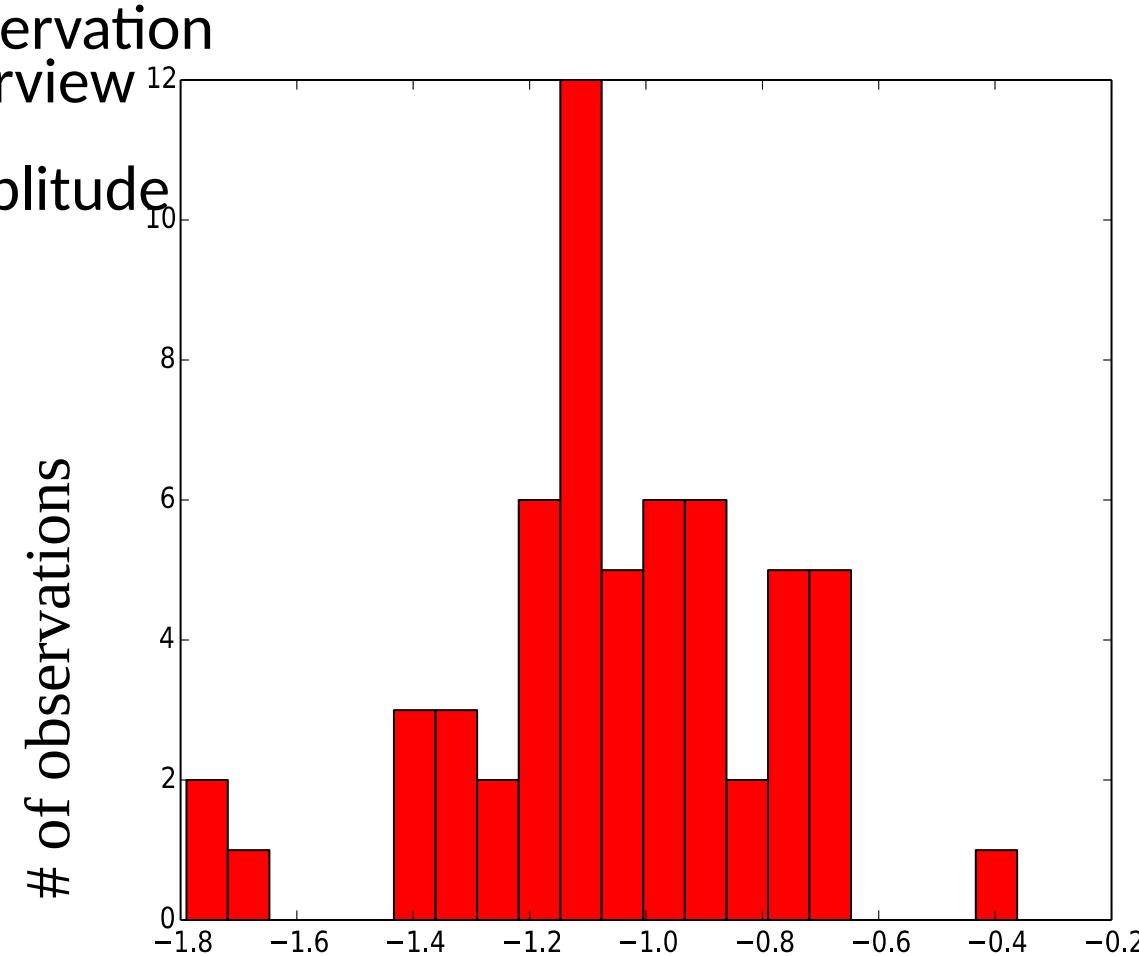
Intro  
Librar  
y

Pulsar  
S  
pin up  
/down

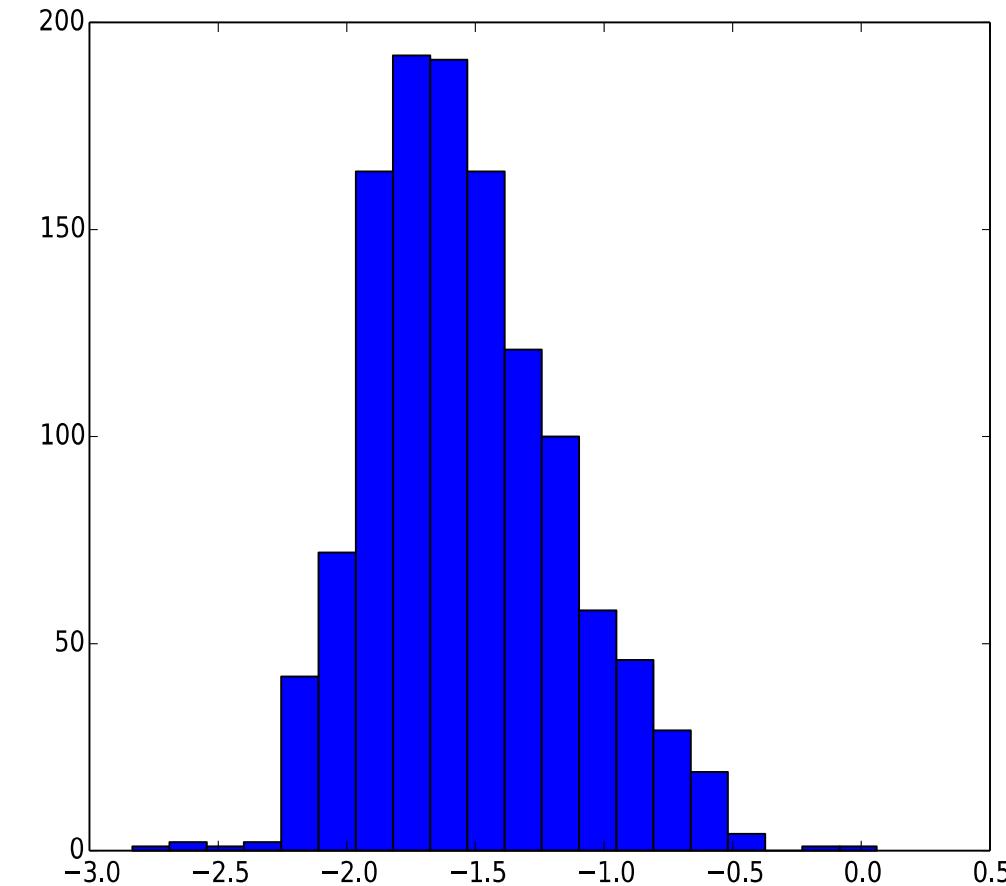
Pulse  
profile  
model

In sum

Observation  
overview



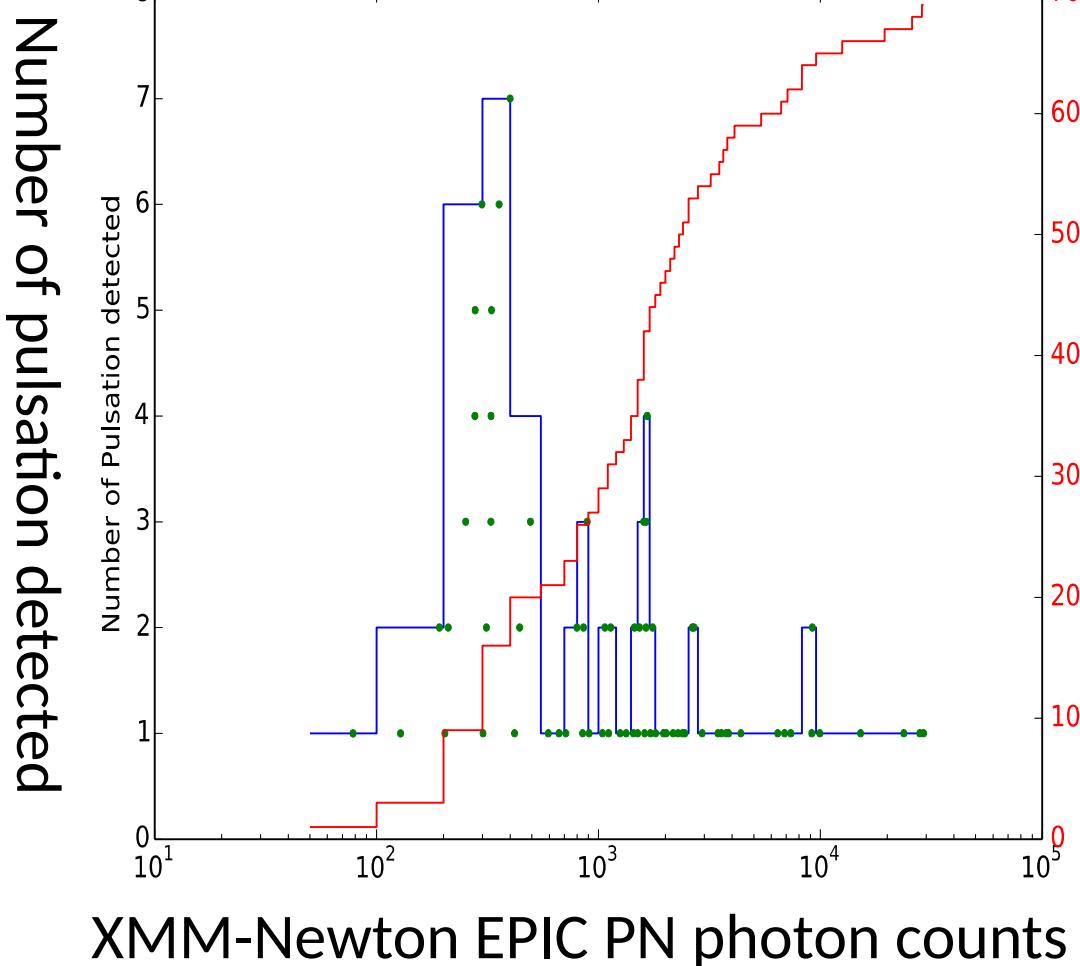
Log (pulse amplitude (cnt/s))  
with Chandra Observations



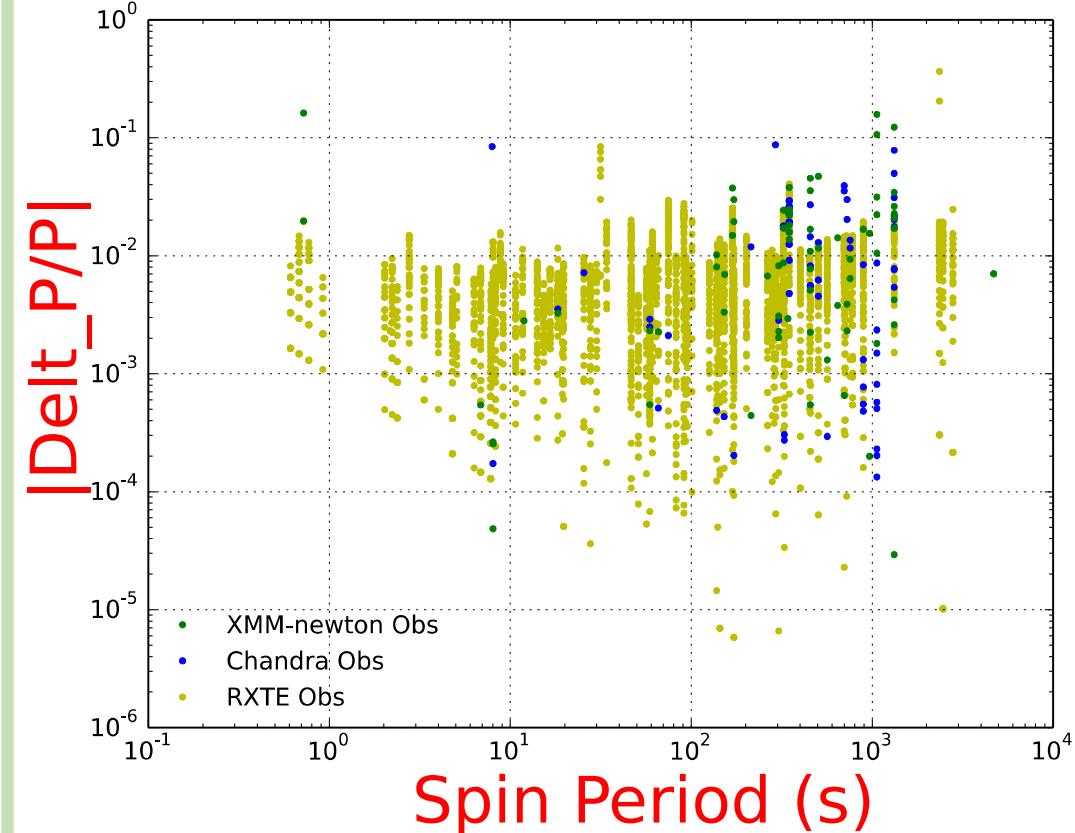
Log (pulse amplitude (cnt/s)) with  
XMM-Newton Observations

# SMC Pulsar library products

Pulsars  
In SMC  
&Mode  
I  
Jun  
Yang  
Intro  
Library  
**Pulsar  
s  
spin  
up  
/down**  
Pulse  
profile  
model  
In sum



Cumulative # of pulsation detected



Period  
deviation

# Library of pulsars in the Magellanic Clouds

Pulsars  
In SMC  
&Mode  
I

Jun  
Yang

**Intro  
Librar  
y**

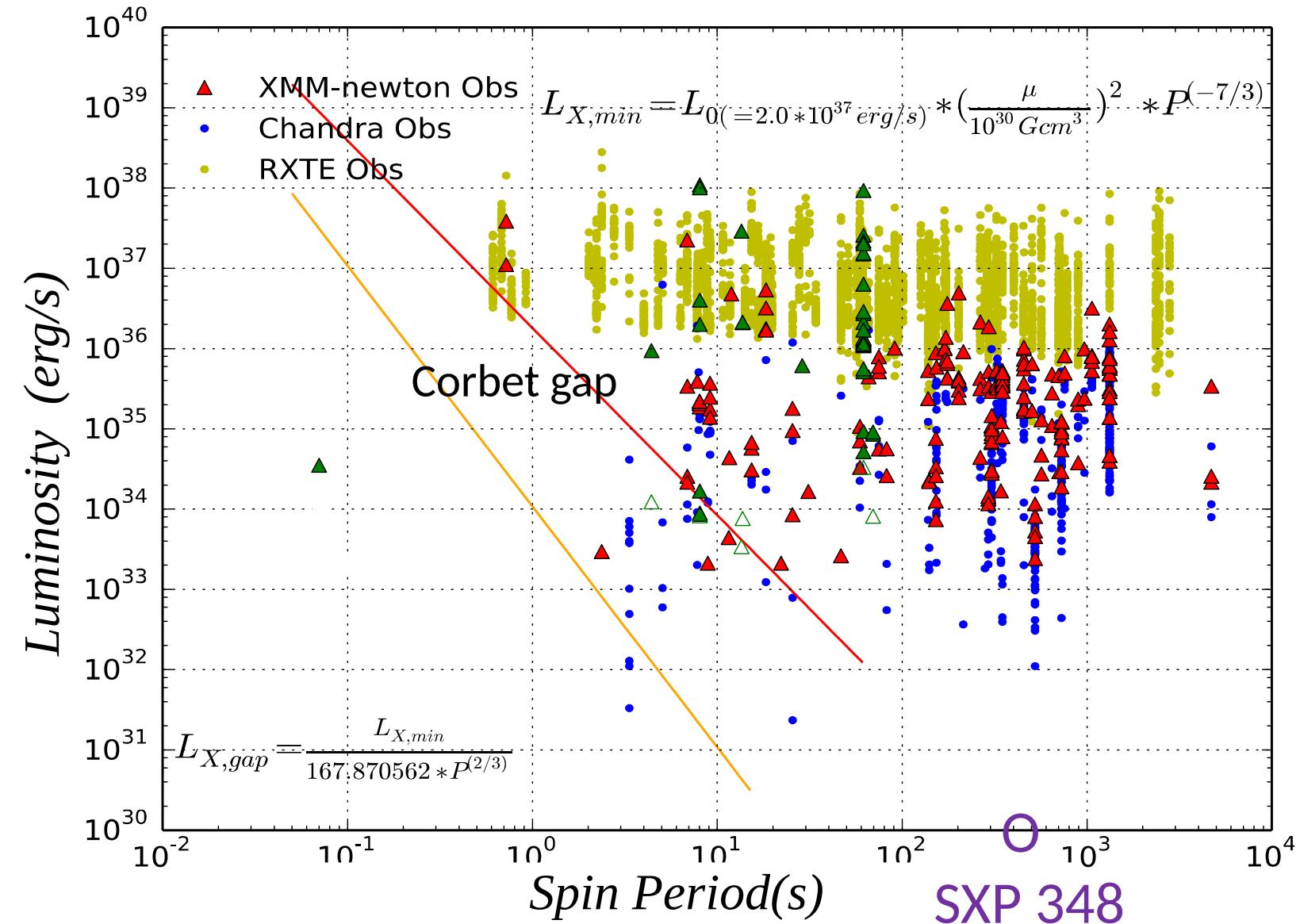
Pulsar  
Ss  
pin up  
/down

Pulse  
profile  
model

## Observation overview

Propeller line  
(P-line) in  
Magellanic  
high mass X-ray  
binaries;

Did not find  
the pulsations  
below the  
P-line



# Makeup

Pulsars  
In SMC  
&Model

Jun  
Yang

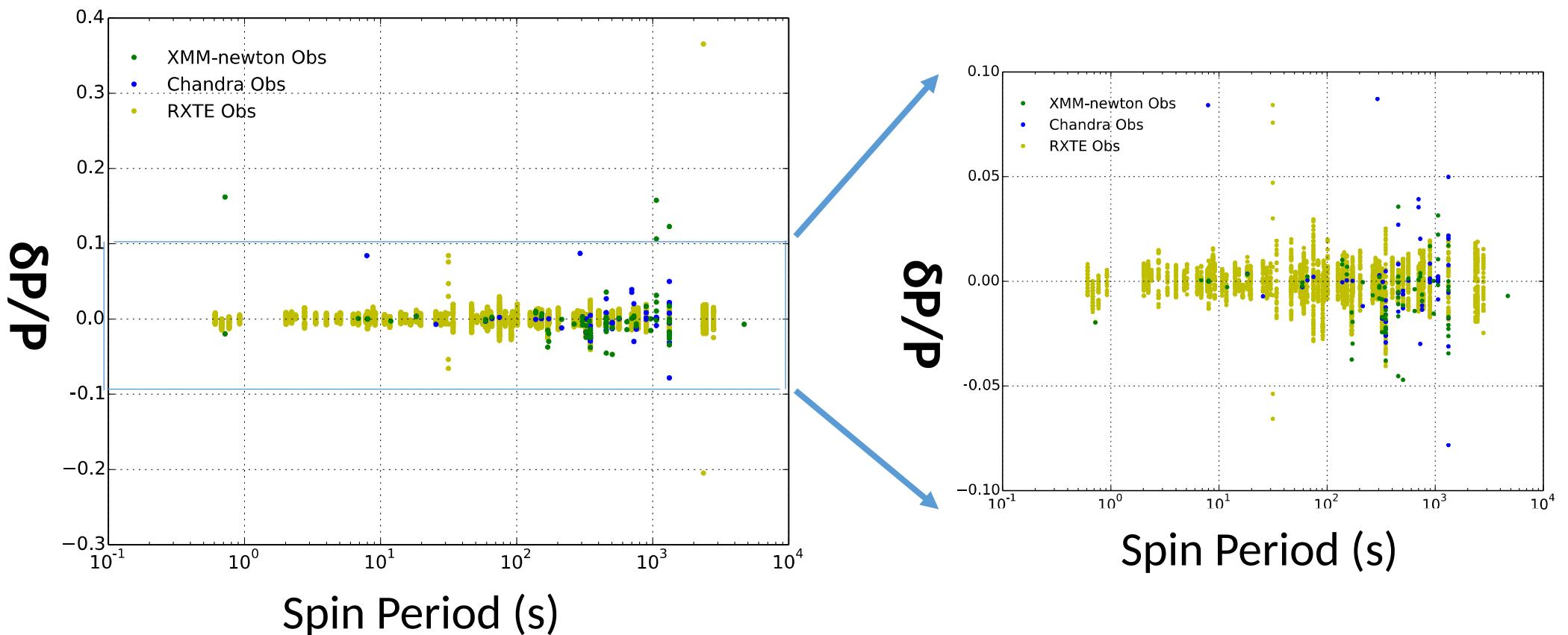
Into  
Library

Pulsars  
spin up  
/down

Pulse  
profile  
model

In sum

**Make  
up**



# Makeup

Pulsars  
In SMC  
&Model

Jun  
Yang

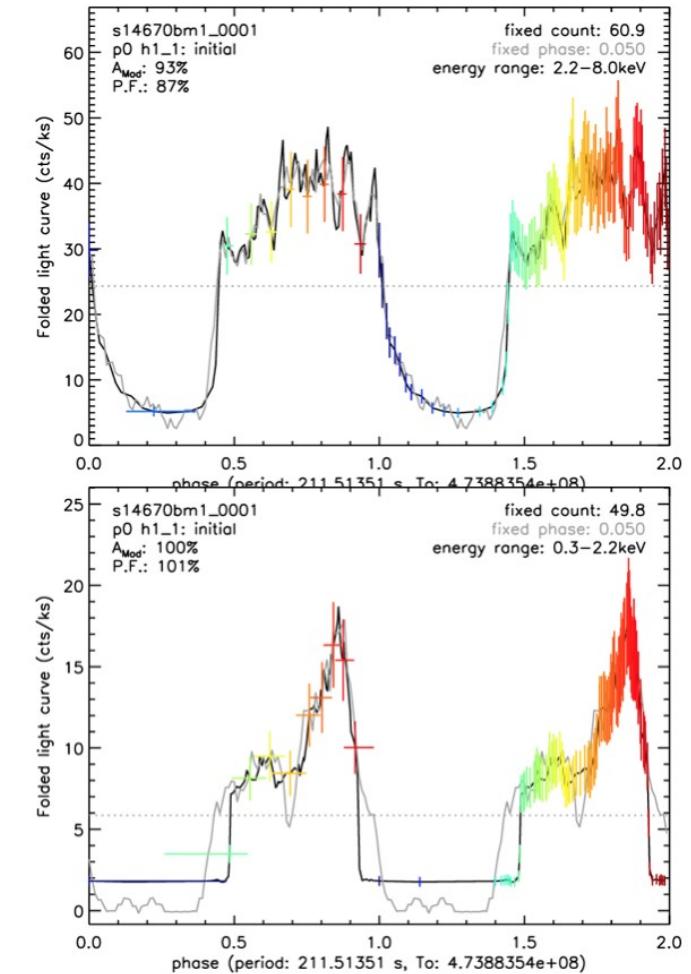
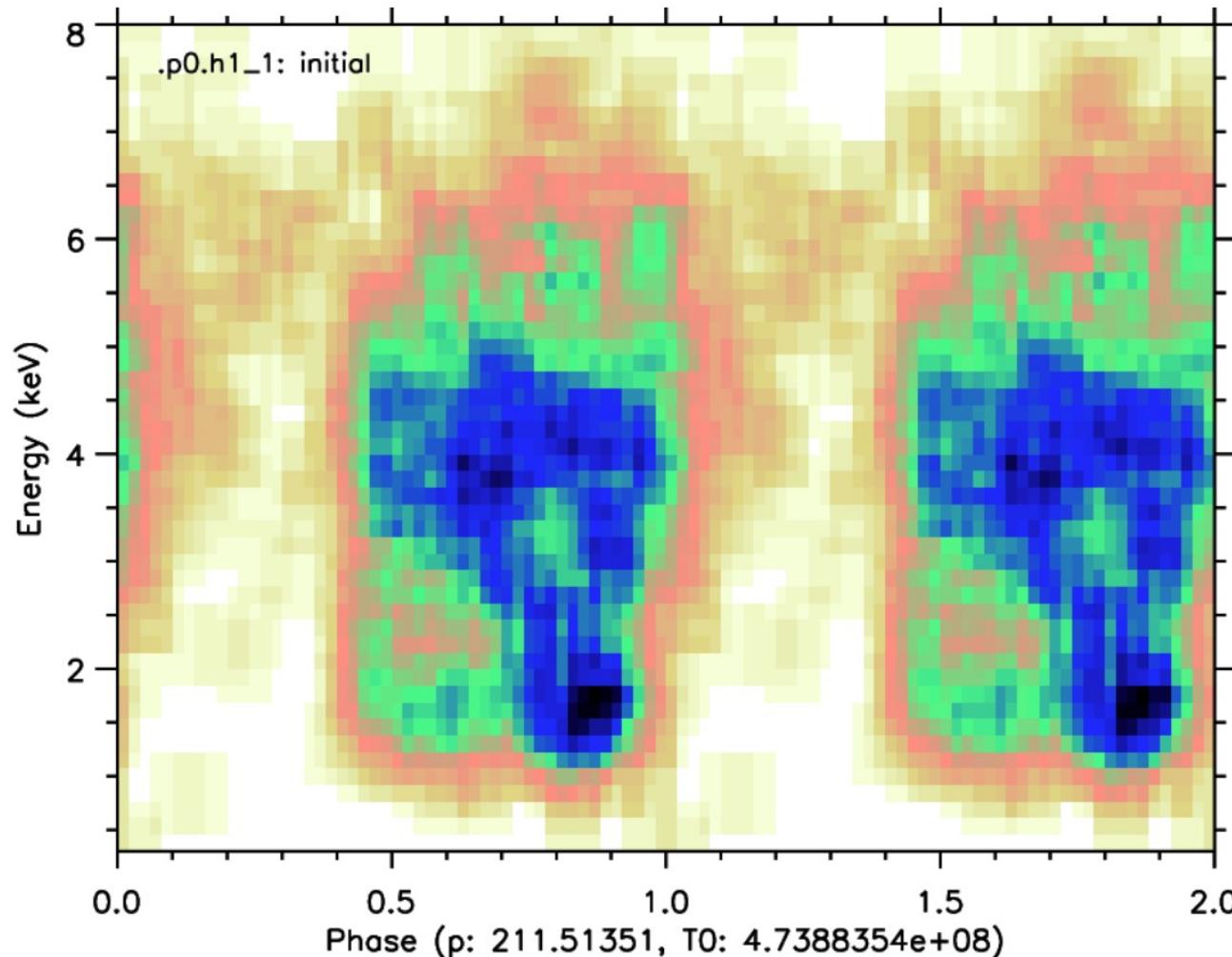
Into  
Library

Pulsars  
spin up  
/down

Pulse  
profile  
model

In sum

**Make  
up**



# Makeup

Pulsars  
In SMC  
&Model

Jun  
Yang

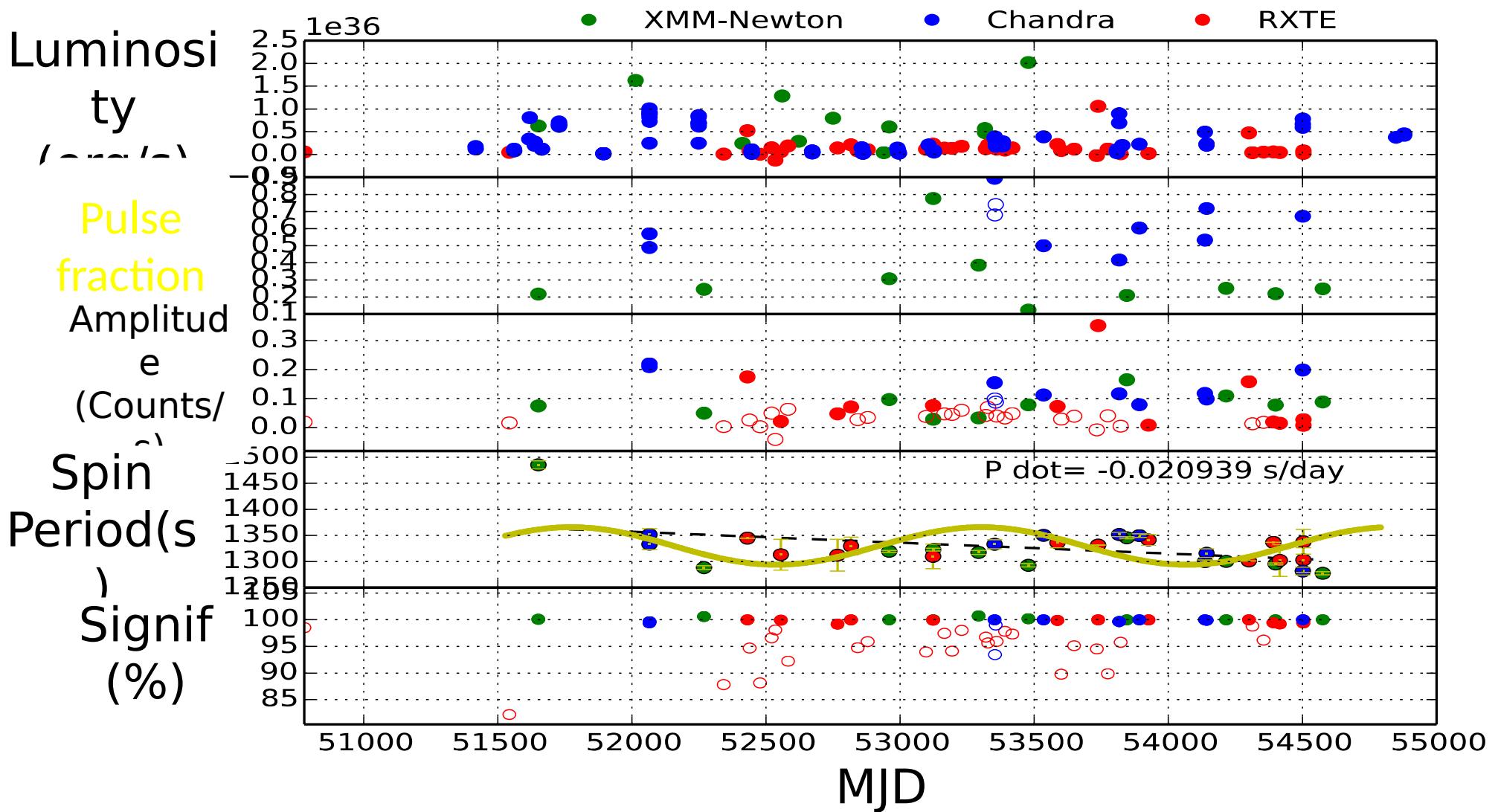
Into  
Library

Pulsars  
spin up  
/down

Pulse  
profile  
model

In sum

Make  
up



SXP 1323 Spin up: 0.0209 s/day

# Makeup

Pulsars  
In SMC  
&Model

Jun  
Yang

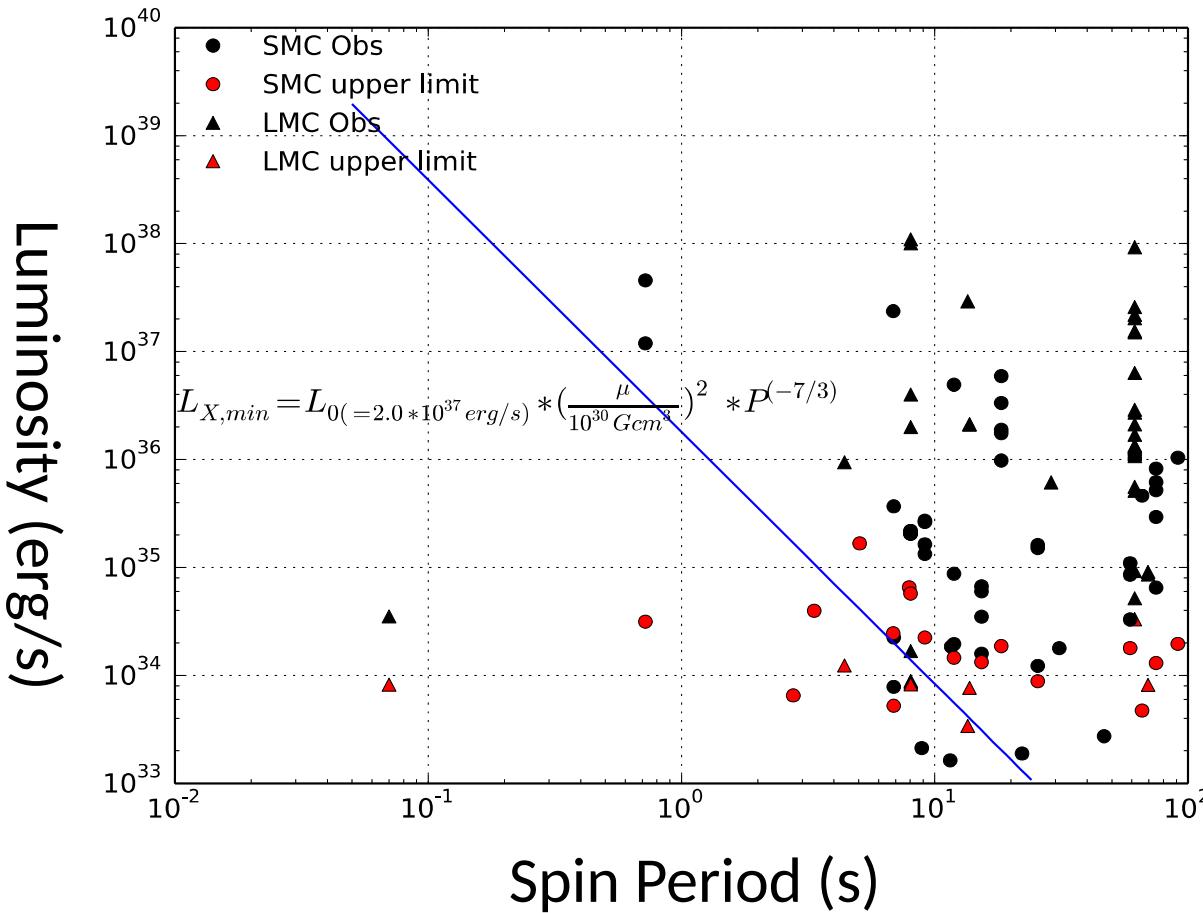
Into  
Library

Pulsars  
spin up  
/down

Pulse  
profile  
model

In sum

Make  
up



# Astrosat

- Broadband Spectroscopic studies of X-ray sources