



Promote AGASC 1.6 for use with OFLS 10.3 and SAUSAGE

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Summary



- **Seek approval to promote AGASC 1.6 for operational use with OFLS 10.3 and SAUSAGE**
- **AGASC 1.6 corrects a calibration error in AGASC 1.5. For red stars the predicted magnitudes are up to 0.5 mags too bright**
- **New version also has improved star magnitude uncertainties which account for stellar spectral variation**
- **Only the MAG_ACA and MAG_ACA_ERR columns have been updated. All other values are identical and files retain the same length and same number of header blocks.**

Presentation:

- **Background**
- **Calibration**
- **Creation of AGASC 1.6**
- **Validation**



Background



- AGASC 1.5 uses observed V and B-V color to predict star magnitude in the ACA bandpass, which is fairly red
- Predicted magnitude can have significant uncertainty because of variations in stellar spectra
- AGASC 1.5 was calibrated in May 2001 using 1939 observed Tycho-1 stars to update MAG_ACA and MAG_ACA_ERR
- Derived a third order polynomial fit

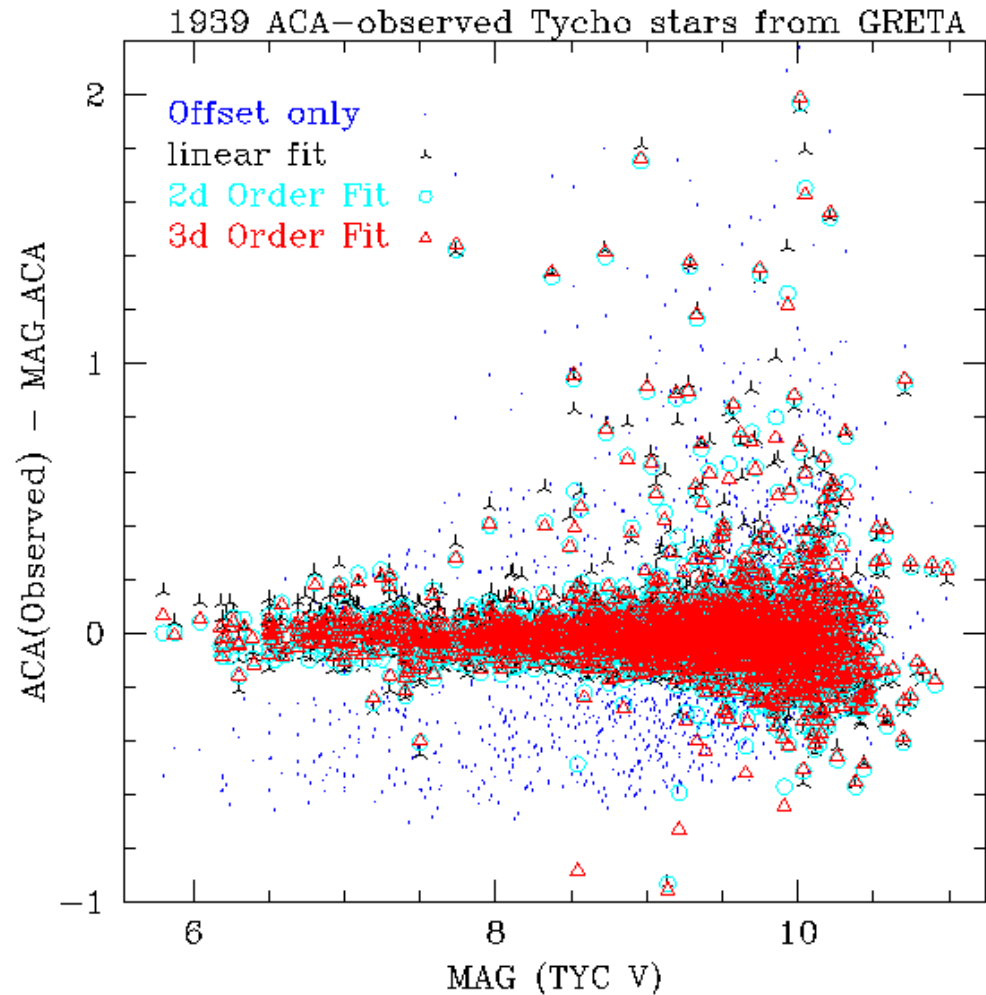
$$\text{MAG_ACA} = V + C_0 + C_1*(B-V) + C_2*(B-V)^2 + C_3*(B-V)^3$$



Background



Apparently a good fit, but...

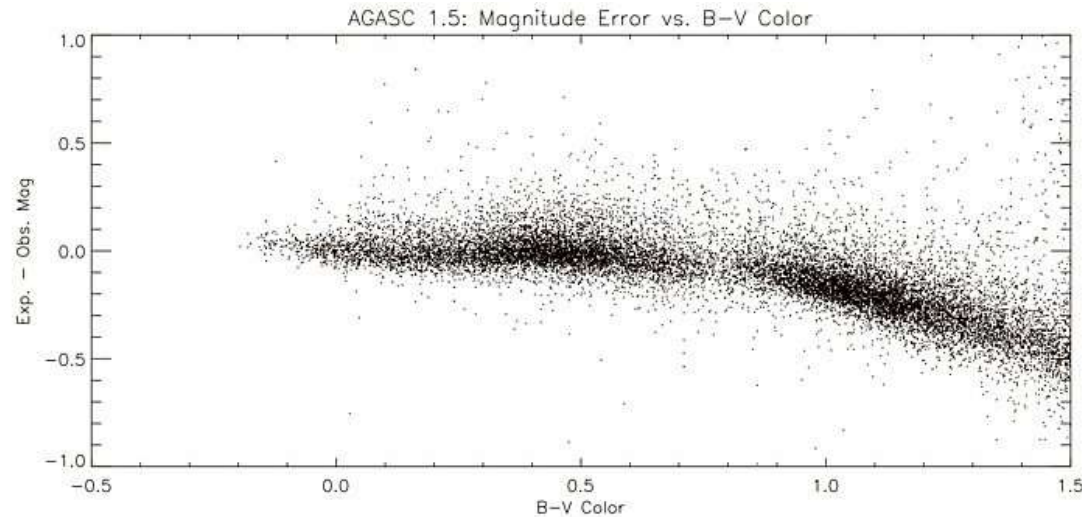


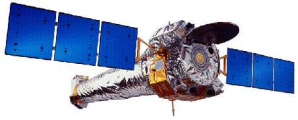


Background



- **Early in 2004 it was observed that for red stars ($B-V > 0.5$) there was an increasing discrepancy between the predicted and observed ACA magnitudes (using 25937 acquisition stars)**

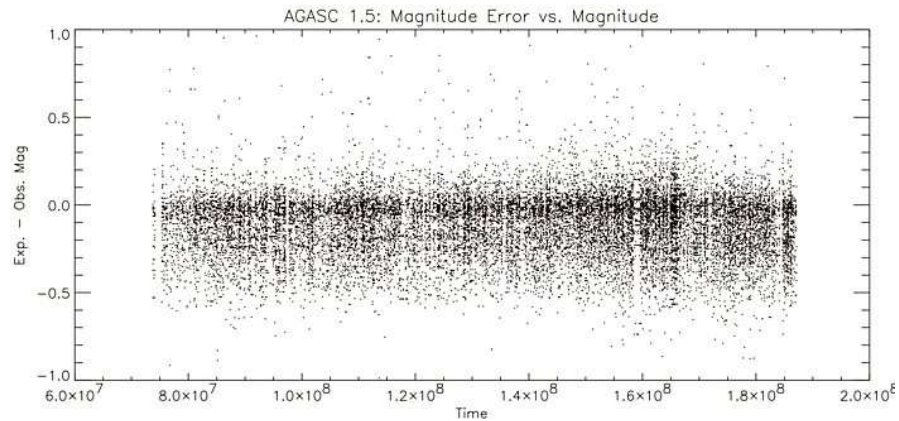
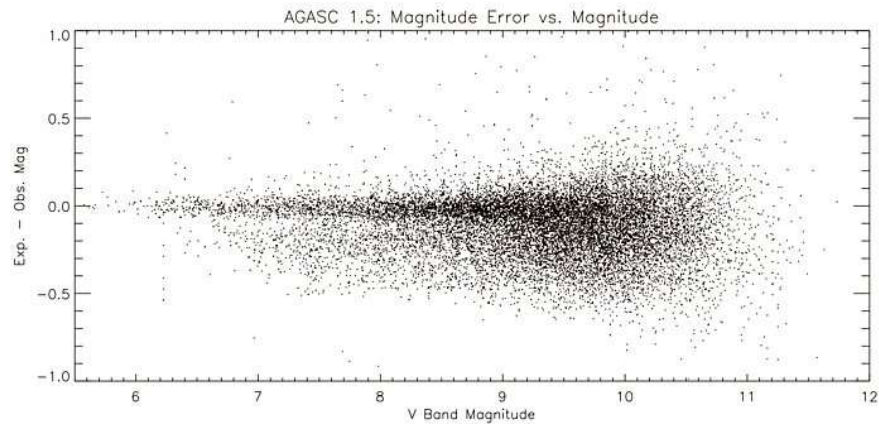




Background



- There is no strong magnitude or temporal dependence





Background



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- The discrepancy between predicted and observed magnitudes was eventually traced to a “accounting” error in program used to generate AGASC 1.5:
 - The usual Johnson B-V color is related to the Tycho $B_T - V_T$ color by $B-V = 0.85 * (B_T - V_T)$
 - In the process of generating AGASC 1.5 that correction factor got applied in twice in different places
 - The star selection and acquisition working group came to a consensus that a fix was needed and that the cleanest legacy solution was to correct the AGASC



Calibration



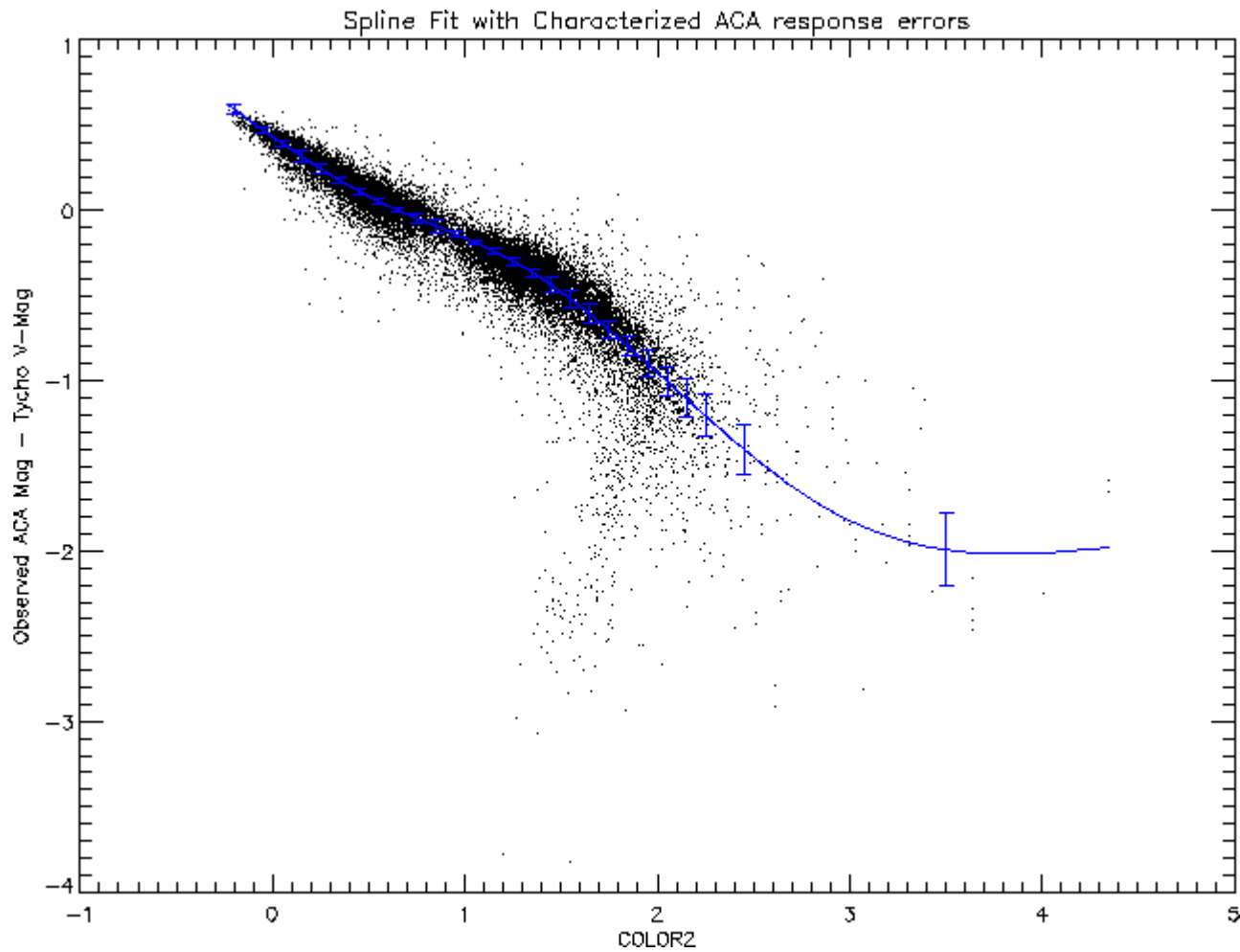
- Using a substantially larger database of observed stars (~26000 vs. 2000), we generated a new best-fit curve to relate MAG_ACA and Tycho-2 magnitudes B_T and V_T .
- Recalibration only applies to stars with Tycho-2 colors (>99.8% of candidate guide/acquisition stars)
- With larger database we could fit over a wider color range
- A spline fit was used instead of a polynomial fit, giving a better fit over the wide range
- MAG_ACA_ERR was recalculated to account for intrinsic uncertainty related to dispersion in stellar spectra

$$\sigma_{\text{mag_aca}}^2 = \sigma_v^2 + \sigma_{(b-v)}^2 [d \text{MAG_ACA} / d(B-V)]^2 + \sigma_{\text{aca_resp}}^2$$

- ACA response term calculated so mean MAG_ACA_ERR matches observed RMS in each color bin
- Details: <http://cxc.harvard.edu/mta/ASPECT/agasc1p6cal/>

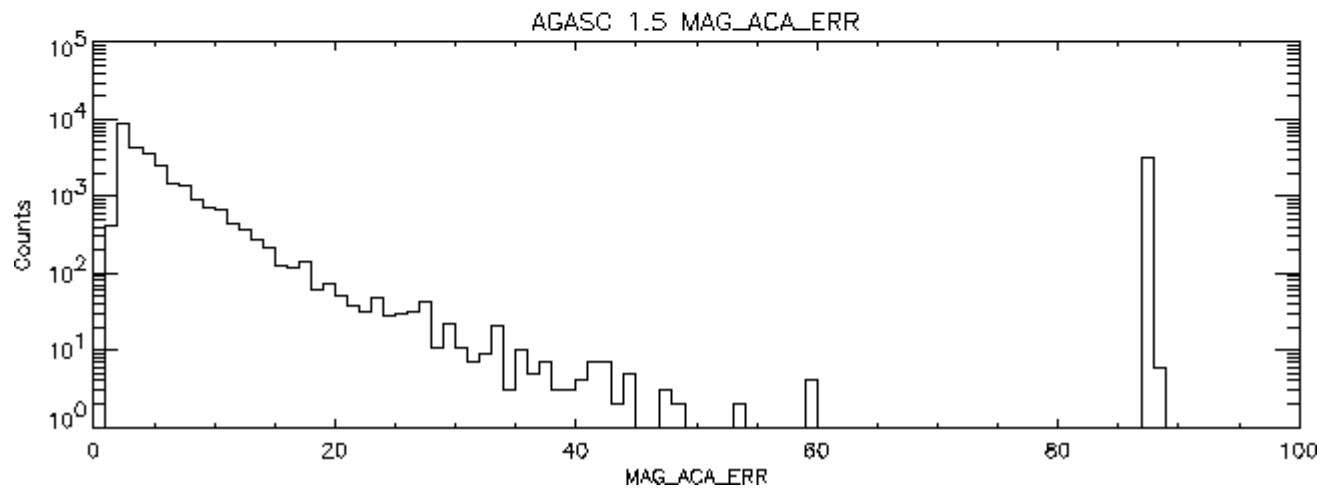
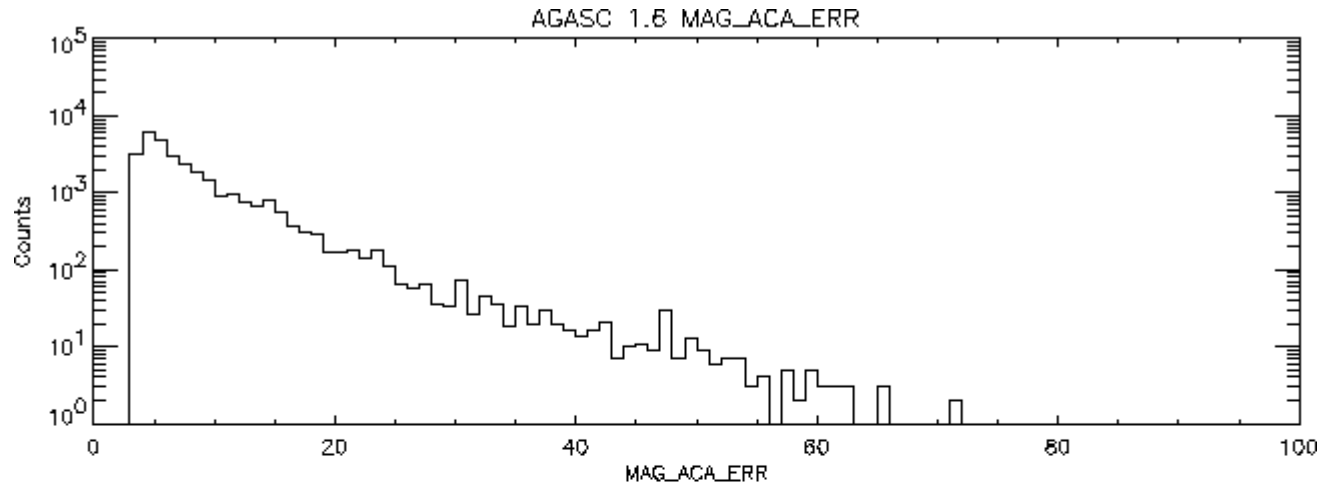


Calibration





Calibration

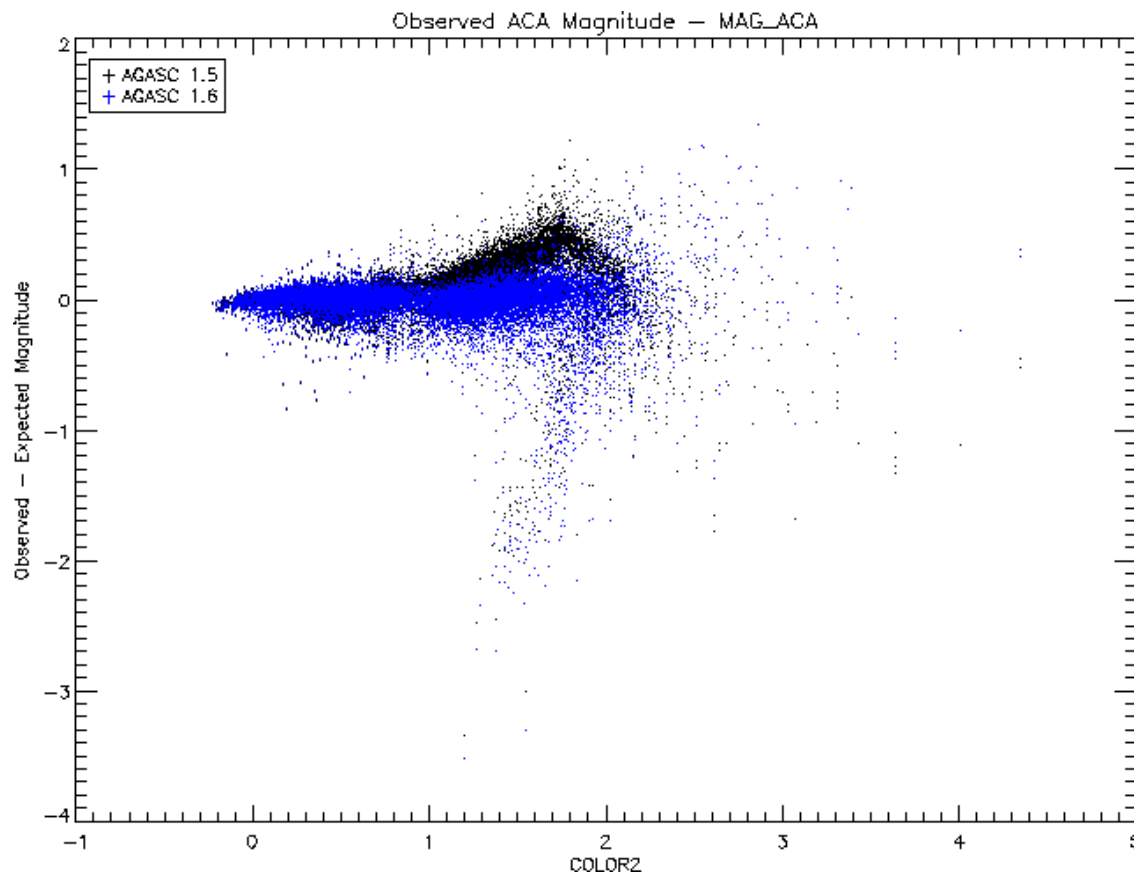




Calibration



- Compare deviations (1.5 vs. 1.6) using actual AGASC 1.6





Creation of AGASC 1.6



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- **The new files for AGASC 1.6 were created using an IDL script**
 - Read in each file
 - Update the **MAG_ACA** and **MAG_ACA_ERR** columns for stars with Tycho-2 colors
 - Update header comments for version 1.6
 - **Documentation**
 - Added AGASC 1.6 entry to main SOT MP AGASC page and supplied links to Aspect web pages with detailed description of calibration and testing
 - Updated standard **COMMENTS.TXT**
 - Put all code and auxiliary data files in a single directory with associated documentation



Validation



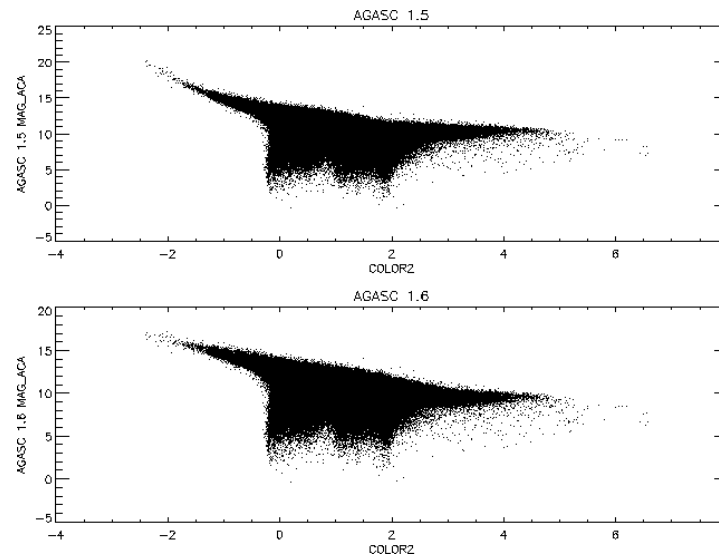
- **Unit level comparison of AGASC 1.5 to 1.6**
<http://cxc.harvard.edu/mta/ASPECT/agasc1p6testing/>
- **CXCDS (Aspect pipeline, MP tools, starcheck, archive)**
- **Flight ops**
 - OFLS 10.3 load product generation
 - SAUSAGE
 - OFLS 10.3 AD&SC
- **Working promotion plan containing detailed test results at**
<http://jeeves.cfa.harvard.edu/Murk/Chandra/PromotionPlan>



Validation: Unit level



- **Unit level comparison of AGASC 1.5 to 1.6**
(<http://cxc.harvard.edu/mta/ASPECT/agasc1p6testing/>)
 - **Color vs. Mag scatter plot**
 - **Color vs. Mag_err scatter plot**
 - **Mag and Mag_err distributions**
 - **Detailed comparison of catalogs generated by SAUSAGE**





Validation: CXCDS



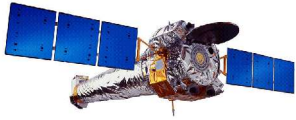
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- **CXCDS Aspect pipeline**
 - Ran aspect pipeline for eight obsids (standard regression test cases) and confirmed no unexpected errors or outputs and only expected changes in guide star output file
 - **Mission planning tools**
 - Confirmed correct operation of `mp_sfe`, `mp_get_agasc_id` and `mp_get_agasc`
 - **Starcheck**
 - Ran starcheck on three FOT test loads and confirmed no errors or unexpected results
 - **SOE file ingest**
 - Ingested SOE file from one FOT test load using beta archive server and confirmed correct ingest



Validation: Flight Ops



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- **OFLS 10.3 and SAUSAGE**
 - AGASC 1.6 installed on SGI/OFLS3 server and NetApp for SAUSAGE access
 - Built three weekly loads and performed full ACA review. (One load specifically chosen to have many red stars)
 - Some differences in selected stars noted, as expected
 - No problems in overall star selection
 - **OFLS AD&SC testing (W. Davis)**
 - Confirmed AGASC 1.6 can be read using W.Davis' PC tool `startest`
 - Ran AD&SC with 14 attitudes with different star densities
 - Results identical for identified stars and attitude solutions
 - **Similar testing will be done with the Linux OFLS 11 at which time we will seek FDB approval for promotion on this platform**



Promotion



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- **Once approved actual deployment will include coordinated events, done at a convenient time between load builds**
 - **ClearDDTS actions:**
 - **Install Release 4 of Star Catalog AGASC 1.6 on OFLS1 in Working State**
 - **Make Star Catalog AGASC 1.6 Operational (Baseline)**
 - **Change star catalog pointer for SAUSAGE**
 - **New DS patch release to change pointers to AGASC 1.6, which is already installed on DS. Precise coordination not required**
 - **After time of no less than 3 months, remove AGASC 1.5**



Conclusion



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- **AGASC 1.6 will improve reliability of star acquisition and guiding**
 - **More accurate predicted star magnitudes**
 - **More realistic assessment of magnitude uncertainties**
 - **Testing shows that there are no radical changes and there will be no impact on the ability to find stars in specific fields**
 - **Important for the legacy of Chandra to have the star catalog correct**