



Science Data Systems

CIAO, Pipelines and Catalog

Jonathan McDowell



CIAO Update - Jonathan McDowell

- Chandra Source Catalog Release 1.0 is out
 - SDS support on analysis, characterization, documentation
 -
- CIAO 4.1 released on schedule in December
- CIAO 4.1.2 patch in April
 - More catalog tools
- CIAO 4.2 release in late 2009
 - More catalog tools
 - First batch of development from CUC input
 - Segmented release (separate Sherpa, Tools, etc)
- Catalog further work
 - Research effort: Coadd overlapping fields
 - Supporting HRC catalog processing.
 - No schedule yet (depends on review of release 1)
- L1 and L2 pipelines:
 - pipeline standard data processing (SDP) is stable; we will respond to any spacecraft issues but no major near-term science changes are expected.
 - “Summary pages” not much used, will stop making them in June 2009



Community support

- CUC input folded into CIAO 4.2 planning
- New computer specialist started Nov 08 – Nick Lee, replaces Nick Durham
 - Big load on computer specialist team during documentation of Catalog and CIAO4.1, all performed outstandingly
- Helpdesk: 90 new tickets (Oct 1 – Mar 18), 11 still open
 - median time to first (useful) answer 1.3 hours, longest time 10d (really a science question; science staff occupied with catalog release)
 - median time to close ticket 11 hours, average 4 days
 - some tickets take longer to close, e.g. waiting for feedback from user, or from internal specialist
 - 45 of 90 tickets handled by helpdesk staff without need for scientist support
- Gave catalog GUI (CSCview) demonstrations at Long Beach AAS meeting



CIAO 4.1

CIAO 4.1 was released in Dec 2008

CIAO 4.1.1 was released in Jan 2009 with updated proposal data files

Downloads since Oct 1 by user's platform: (total 731)

Linux CIAO4.1 186, CIAO4 113, CIAO3.4 20

- Most common: FC8, FC4, Debian

- Still a few FC2,3

Solaris CIAO4.1 14, CIAO4 5, CIAO3.4 0

Mac CIAO4.1 166, CIAO4 82, CIAO3.4 17

- 184 OS 10.5, 81 OS 10.4

Unknown CIAO4.1 75, CIAO4 42, CIAO3.4 9

Current platforms are

Fedora Core 4

Fedora Core 8

Mac PPC OS 10.4

Mac Intel OS 10.4, 10.5

Solaris 8

CIAO 4.1.2 will upgrade Solaris to 10, Mac PPC to 10.5,
evaluate newer Fedora for CIAO 4.2 (successfully using
CIAO on Fedora Core 10 with SELinux workaround)



CIAO 4.1

CIAO 4.1 included the first batch of Catalog-derived tools

Reminder of CIAO 4.1 content (from last CUC):

- **Sherpa**
 - Release 1.0 of the new Sherpa, improved convergence and robustness, support for grating data, user models
- **Science Tools**
 - `aprates` Aperture photometry with confidence intervals (uses user-supplied PSF fraction)
 - `eff2evt` Calculates `eff.area` and 'flux' for each photon in event list
 - `lc_clean` Revised to use new ChIPS
 - `mkpsfmap` Gives image of approximate PSF size for use in detect apps (not for use in aperture corrections)
 - `lim_sens` Limiting sensitivity calculation (for catalog)
 - `dmellipse` Find ellipse containing given fraction of source
 - `dmimgadapt` New adaptive binning tool
 - Upgrades to `wavedetect`, `acis_build_badpix`
- **Infrastructure**
 - DataModel Ascii Kernel - cleaned up, no longer beta - "Release 1" status
 - Image handling improved, can bin tables to images
 - String handling in tables improved, fixed bugs with internal quotes and escape chars.
 - CALDB - new generation CALDB allows better internal organization
 - Prism file viewer improved
 - new interface, better functionality and connection to plots
 - Array columns expanded inline, not in popup; multiple files appear in tabs
 - GUIs (Prism, PEG, TaskMon) - changed out MOTIF for GNU compliant GTK
 - `ardlib` tools now accept data model filters



CIAO 4.1 Sherpa

- Python and Slang environments
 - command line and script/batch mode
 - accessing and working with internal Data Structures
- Improved convergence and robustness with three optimization methods:
 - Levmar, Neldermead and Moncar methods
- Several statistics options
- Support for grating data
- Flexibility in adding user models and statistics functions
- Extensibility of the system for advanced users. Examples include:
 - Simulations to evaluate the confidence on derived parameters such as flux.
 - “Project” models that were developed and posted on the contrib pages by Tom Aldcroft:
<http://cxc.harvard.edu/contrib/deproject>
- Improved on line documentation includes many new threads - it is ongoing



CIAO 4.1.2

CIAO 4.1.2 development is complete and SDS science testing complete:
Completing regression testing and preparing for download testing
Release scheduled for Apr 15
CIAO 4.1.2 contains a set of Level 3 tools, plus bug fixes.

- Sherpa
 - CIAO 4.1.2 Sherpa updated to the latest XSPEC12.5 models
- Science Tools
 - **evalpos** Simple script to get image pixel value at a celestial position (e.g. for sensitivity file values)
 - **glvary** - Variability tool developed for catalog
 - **modelflux** – New tool based on catalog script, to convert between flux and rate for arbitrary Sherpa model
 - **srcextent** – Modified from catalog tool for better user interaction; measures effective size of source and PSF and estimates intrinsic size.
 - **streakmap** - Modified from catalog tool to accept more general user inputs; calculates streak background map due to CCD readout 'out-of-time' events.
 - **pileup_map** - Simple script to estimate pileup based on frame count rates.



CIAO 4.1.2

TESTING CIAO 4.1.2

CIAO 4.1.2 is undergoing science testing and regression testing by SDS.

For each tool the assigned scientist prepares a worksheet giving several test cases and results. The test team uses them to prepare scripts for portable regression testing.

CIAO 4.1.2 SDS carries out regression testing on multiple platforms. We have upgraded our testing scripts and machines, and we now test on

- Fedora Core 8,
- Mac OS 10.4 PPC,
- Mac OS 10.4 Intel,
- Solaris 8,

and are beginning tests on Solaris 10. For Mac Intel testing we have been borrowing a scientist's desktop, but we are now getting a dedicated machine whose configuration can be more easily controlled.



CIAO 4.1.2 Gratings Documentation



New: Grating Analysis Guide
Links together existing threads for ACIS/HETG and ACIS/LETG work

New Line fitting thread

ANALYSIS GUIDES (for HETG/ACIS-S, LETG/ACIS-S, LETG/HRC-S) <http://cxc.harvard.edu/ciao/guides/>

- Roadmap of the CIAO threads for grating data analysis (high resolution spectroscopy)
- Takes the user through the analysis of a single source from a single observation.

There is also a: *High Resolution Spectroscopy - Detailed Overview* (by Dave Huenemoerder)
good comprehensive overview of grating processing and analysis (old, needs some updating).



CIAO 4.1.2 Gratings Documentation



New: Grating threads for Sherpa

CIAO/SHERPA DATA ANALYSIS THREADS:

Grating spectroscopy section: <http://cxc.harvard.edu/ciao/threads/gspec.html>

Include about 30 threads for all grating combinations (including LETG/HRC-I!) that explain how to examine pha2 files, create color spectra, correct zero-order source position, create grating ARFs and RMFs, group and coadd spectra.

Modeling & Fitting Grating Data with Sherpa: <http://cxc.harvard.edu/sherpa/threads/fitting.html>

Fitting Grating Data: General introduction to fitting grating data in Sherpa. Loading and filtering data are covered, as well as defining instrument responses and source models.

Fitting Multiple Orders of HRC-S/LETG Data: Explains how to model overlapping orders in Sherpa by defining the instrument response as a composite of the interesting orders. The thread uses response files (gRMFs and gARFs) created with CIAO to model and fit the first three +/- orders of the spectra.

Measuring Line Parameters: After you have a set of PHA2 and response files from TGCAT (for an HETG/ACIS-S observation), perform a simple fit to one of the line features present in the spectrum.

Simulating an ACIS/HETG emission-line grating data set: using responses and backgrounds from an observed data set

TGCAT SOFTWARE MANUAL: http://space.mit.edu/cxc/analysis/tgcat/tgcat_man.pdf

Explains how to run the reprocessing scripts to obtain pha2 files, grating ARFs and RMFs, light curves, and quick-look summary plots



TGCat

The gratings team at MIT has reprocessed HETG data and released products at

<http://tgcat.mit.edu>



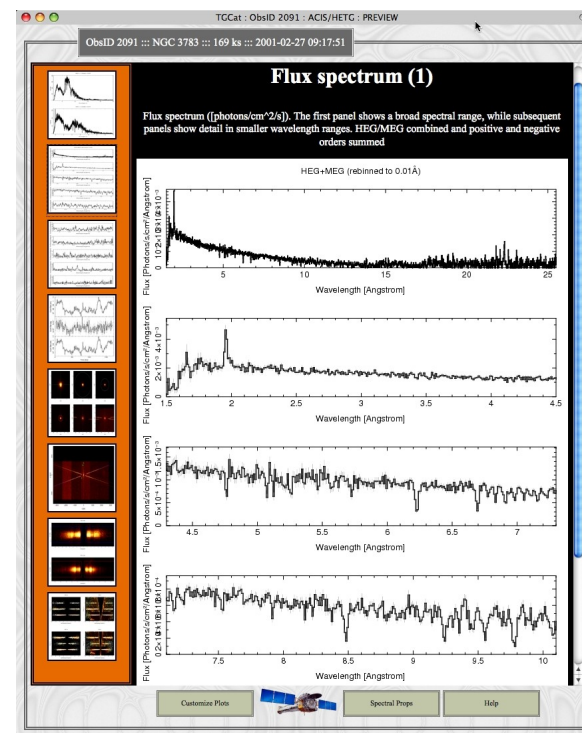
RESULTS: Found 152 matching extractions

ObsID	Count	Source	Instrument	Grating	RA (h:m:s)	Dec (d:m:s)	Date Obs (y-m-d)	Exposure (s)
1673	2482	NGC 3516	ACIS	HETG	11:06:47.467	+72:34:7.176	2001-11-11 01:00:25	88011.4
1673	8451	NGC 3516	ACIS	HETG	11:06:47.484	+72:34:7.212	2006-10-11 09:49:34	47567
1673	7281	NGC 3516	ACIS	HETG	11:06:47.501	+72:34:7.284	2006-10-14 02:19:19	42509.5
1673	2080	NGC 3				+72:34:7.248	2001-04-10 17:55:54	73376.6
1673	8452	NGC 3				+72:34:7.176	2006-10-09 14:05:36	20754.8
1673	2431	NGC 3				+72:34:7.320	2001-04-09 14:12:05	35572
1673	7282	NGC 3				+72:34:7.212	2006-10-10 04:43:54	41454.5
1673	831	NGC 3				+72:34:7.104	2000-09-30 21:05:22	43893.8
1673	8450	NGC 3				+72:34:7.248	2006-10-12 07:00:26	38549.6
1666	2092	NGC 3783	ACIS	HETG	11:39:1.694	-37:44:18.960	2001-03-10 00:31:15	165454

Flux Spectrum (Click to preview all Images)

press "go" to operate on selections: limit download plot combined view sources

Go Change Columns New Search Help





CIAO 4.2

CIAO 4.2 (Dec 2009) planning well underway

Focus is point source analysis, and more catalog-derived tools
Below I list examples of SDS development work for 2009

- Candidate Science Tools for CIAO 4.2 - catalog
 - BEHR BEHR algorithm for hardness ratios
 - flux_eval Use eff2evt to make flux estimate, add confidence intervals
 - background map tool
 - error ellipse combination tool
- Candidate Science Tools for CIAO 4.2 – enhancing existing tools
 - acis_build_badpix Tool to make bad pixel files for your observation, adding user-requested improved abilities to customize
 - acis_process_events support for graded CTI [request from Cal team]
 - acis_process_events support for sub-pixel resolution
 - hrc_process_events support for new gain files
 - sherpa Enhance user interface
- R&D work
 - MARX/ChaRT comparison and characterization; PSF aperture correction for ARF



CIAO 4.2

CIAO 4.2 scripts, GUIs and threads

- Restore ChIPS plotting functionality to user scripts (lost in CIAO3.4-4.0 upgrade) and improve interface
- Work on scripts for merging observations
- Add statistical error analysis scripts based on CHASC astrostatistics collaboration
- Simple user data-reprocessing scripts
- Investigate simple GUIs for Sherpa (long lead)



CIAO 4.2 and beyond: user interface

- We are currently supporting two advanced scripting languages in Sherpa/Chips: S-lang and python
- This is a major support headache – the documentation is OK (mostly automated) but science testing of two versions of each function is too much work
- python is working well in our system after an initial learning curve
- Following CUC recommendation in Sep 07, we are considering phasing out S-lang in the Sherpa/Chips environment and reviewing the status of S-lang modules in CIAO
- S-lang would still be packaged with CIAO and users can continue scripting in it outside of Sherpa/Chips
- However, we are concerned about impact to existing users – welcome CUC feedback. Will keep CUC informed as we evaluate options.
- Also considering additional user interface work following feedback from some users who prefer a command line interface and find the functional style confusing.



Catalog publications

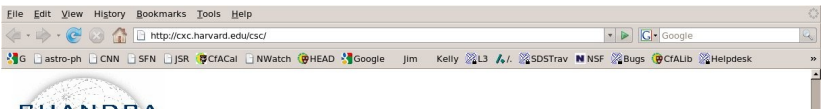
- SDS catalog posters presented at Long Beach AAS:
 - Chandra Source Catalog: Algorithms (McDowell et al)
 - Chandra Source Catalog: Source Variability (Nowak et al)
 - Chandra Source Catalog: X-ray Aperture Photometry (Kashyap, Primini et al)
 - Chandra Source Catalog: Statistical Characterization (Primini et al)
 - Chandra Source Catalog: User Interface (Bonaventura et al)

in addition to posters authored by software team (I Evans et al, overview; J Evans et al, processing; S Doe et al, spectra; R Hain et al, merging)

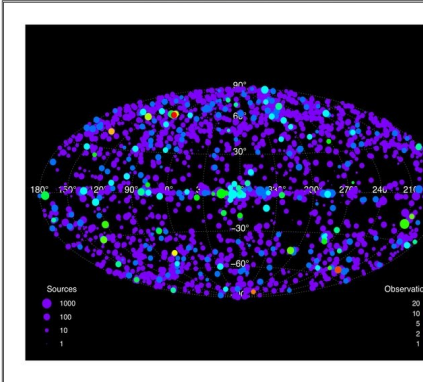
and by archive team (A Rots et al, source properties; M McCollough et al, backgrounds)



Catalog documentation

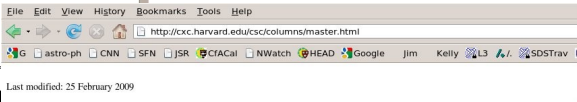


The Chandra Source Catalog Release 1.0: Point and compact source catalog



The locations of observations included in the CSC. In Galactic coordinates (click the image for equatorial). The size of each symbol is proportional to the logarithm of the number of sources detected in the field while the color encodes the number of closely-fleeting observations.

- CSC Data Access (CSCview)
- CSCview Software Requirements
- CSCview Help
- CSC Homepage
- About the Catalog
- Catalog Organization
- Catalog Release Views and Database Access Views
- Catalog Statistical Characterization
- Schedule and Status
- Caveats and Limitations
- Creating the Catalog
- Observation Selection
- Catalog Processing
- Data Products
- Chandra Data Archive
- Using the Catalog
- Threads
- Level-3 Files
- CSCview GUI
- Catalog Columns
- Master Chandra Source Table: alphabetical | by context
- Table of Individual Source Observations: alphabetical | by context
- Column Descriptions
- Position and Position Errors
- Source Flags
- Source Extent and Errors
- Energy Bands
- Source Fluxes
- Source Significance
- Spectral Properties
- Source Variability



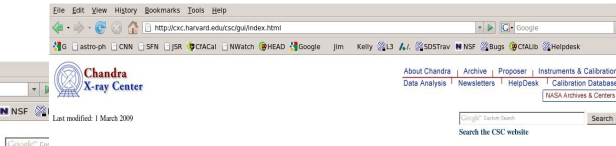
Master Chandra Source Table

Each identified distinct X-ray source on the sky is represented in the catalog by a single "master source" entries, one for each observation in which the source has been detected. The master source entry records the source, based on the data extracted from the set of observations in which the source has been detected.

Go to: [Catalog Columns Index](#) | [Alphabetical List](#)

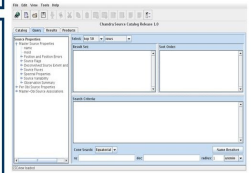
Context	Column Name	Type	Units
Source Name	name	string	
Position	ra	double	deg
Position	dec	double	deg
Errors	err_ellipse_r0	double	arcseconds
Errors	err_ellipse_r1	double	arcseconds
Errors	err_ellipse_ang	double	deg

- CSC Data Access (CSCview)
- CSCview Software Requirements
- CSCview Help
- CSC Homepage
- About the Catalog
- Catalog Organization
- Catalog Release Views and Database Access Views
- Catalog Statistical Characterization
- Schedule and Status
- Caveats and Limitations
- Creating the Catalog
- Observation Selection
- Catalog Processing
- Data Products
- Chandra Data Archive
- Using the Catalog
- Threads
- Level-3 Files



Using CSCview

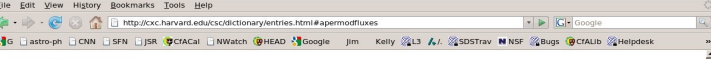
Introduction



Threads

- Conducting a Property Search
- Conducting a Cross Search
- Retrieving Data Products
- Using Source Property Associations

The Chandra Source Catalog (CSC) is ultimately intended to be the definitive catalog. To achieve that goal, the catalog will be released to the user community in a series of releases of the CSC includes information about sources detected in public ACI Chandra mission, only point sources, and compact sources, with observed extended sources, and sources located in selected fields containing bright, high



CSC Dictionary Entries

Alternating Exposure Mode

In alternating exposure mode, sometimes also referred to as interleaved mode, there is a long (primary) and a short (secondary) frame time, and the observation alternates between them. The long frames are joined into one event file and the short frames are joined into another event file. This mode is used to look for variability on different timescales or when observing a source that is expected to be piled. The observer gets the long-frame "piled" information, as well as the short-frame unpiled data.

Aperture Model Energy Fluxes

The conversion from aperture source count rates in each science energy band to aperture model energy fluxes is performed by scaling from a model spectrum folded through the calibrated response, as follows:

For a source model $F(E)$ whose integral over the science band is $F(\text{band})$, calculate the corresponding band count rate $C(\text{band})$ in counts s^{-1} , given the effective area calibration $A(E)$ (and, if available, the RMF_i , $R(E)$, $E(\text{band})$) appropriate to the observation; this is the integral of $F(E)A(E)/R(E)$ over all energies, or if a diagonal RMF is assumed, the integral of $F(E)A(E)$ over the band.

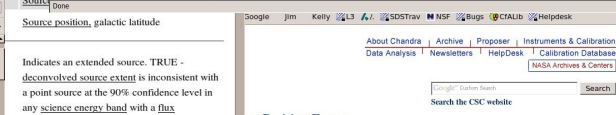
Infer the aperture model energy flux from the measured aperture source count rate $C(\text{band})$ as $F(\text{band}) = F(\text{band})C(\text{band})/C(\text{band})$. The power law spectral model has a fixed photon index, α , defined as $F_{\alpha} e^{-E/\alpha}$, equal to 1.7, and a fixed total neutral Hydrogen absorbing column $N_H = N_H(\text{Gal}) \text{ cm}^{-2}$. The black body spectral model has a fixed temperature $kT = 1.0 \text{ keV}$, and a fixed total neutral Hydrogen absorbing column $N_H = 3 \times 10^{20} \text{ cm}^{-2}$.

Return to: [Dictionary index](#)

Aperture Total Counts

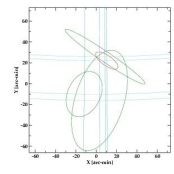
"Aperture total counts" refer to the total number of source plus background counts measured in the modified source region, the modified background region, the modified elliptical aperture, and the modified elliptical background aperture, uncorrected by the PSF aperture fraction.

Return to: [Dictionary index](#)



Position Errors for Chandra Source Table

1D and 2-dimensional positional uncertainty of a source listed in the Master Chandra Source Table represents the best estimate of the source position based on several independent entries. The master source entry is the merged result of multiple observations of the same source. To determine the best estimate of the position of a source from several estimates of its position, we employ a 2-dimensional optimal weighting formalism to statistically average the source positions resulting from the set of individual observations. It is decided to use this technique because it offers an improved estimate of source position where simple averaging fails, e.g. where the area defining the source positions varies from measure to measure. We express the uncertainties of the estimates in the form of error ellipses centered upon the estimated source positions.



An example of input ellipses (green) and the combined ellipse (blue). The x and y values represent $\ln(\text{flux})$ coordinates.

These that are combined to produce the "best estimate" error ellipse for the merged source entry result from the Chandra Multiwavelength Project (ChMP) positional uncertainty described in the page "Source Position Errors in the Table of Individual Source Observations."

The multivariate optimal weighting formalism in the Chandra Source Catalog, described below, represents its first application to astrophysical data, as it is based on an analysis of data from a Master's thesis from the Naval Postgraduate School. For more information on this analysis and its use in the Chandra Source Catalog, see Joseph R. Orosco's "Single Source Error Ellipse Combination" and John Davis' CSC document "Combining Error Ellipses," respectively.

Main page
column descriptions
GUI usage
dictionary
special topics
etc...

CUC meeting Apr 2009



Catalog – Release 1.0 support

Catalog release process involved significant work by SDS team in support of DS software release effort

Documentation updated

Example science thread for users

User testing of CSCView GUI interface

Web site infrastructure and release

Science support for last-minute data migrations following discovery of minor bugs in the data found by science characterization review